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Non-equilibrium dynamics of a nonlinear Jaynes–Cummings model in cavity arrays

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Abstract

We analyze in detail an open cavity array using mean-field description, where each cavity field is coupled to a number of three-level atoms. Such a system is highly tunable and can be described by a Jaynes–Cummings like Hamiltonian with additional nonlinear terms. In the single cavity case we provide simple analytic solutions and show, that the system features a bistable region. The extra nonlinear term gives rise to a rich dynamical behavior including occurrence of limit cycles through Hopf bifurcations. In the limit of large nonlinearity, the system exhibits an Ising like phase transition as the coupling between light and matter is varied. We then discuss how these results extend to the two-dimensional case.

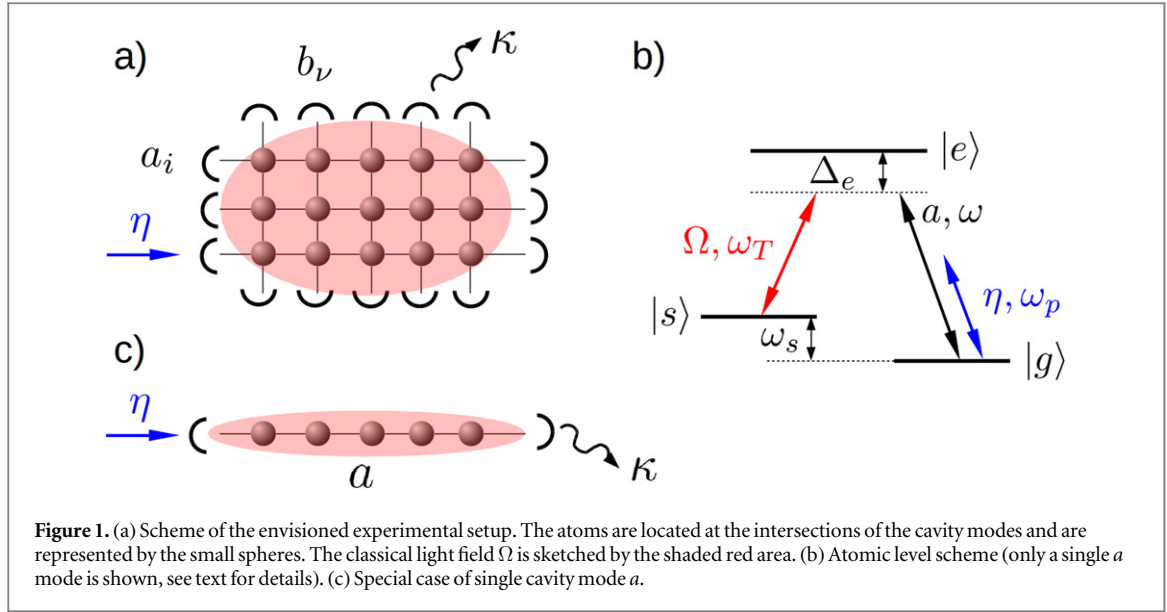
1. Introduction

The use of cavity QED tools is now ubiquitous across different areas of physics ranging from quantum information [1, p 75] to detection of dark matter [2]. Specifically, the atoms held in optical cavities play a vital role in studies of many-body physics [3]. Such systems are natural implementations of many-body Jaynes–Cummings (JC) or Dicke Hamiltonians [4]. Their high tunability and the possibility of achieving strong light matter coupling or probing the dynamics in real time make them very attractive experimental platforms. The prospect of probing phase transitions and the associated critical phenomena with these platforms have been put forward e.g. in [5–7] and the non-equilibrium dynamics of the Dicke model has been theoretically investigated in [8]. Specific many-body phenomena that can be studied with cavity QED include for example the physics of spin glasses [9, 10], the emergence of gauge fields and the related quantum Hall effect [11] or the self-organization of the atomic motion [12–14], to name a few. The self-organization has been subsequently observed in the experiments [15, 16].

So far we have mentioned only studies concerning a single cavity—generalizations to multiple cavity arrays implementing the Hubbard physics have been reviewed in [17]. Although appealing in principle, the realization of many efficiently coupled cavities, each hosting a discrete-level quantum system is a challenging task. To make such experiment scalable requires miniaturization of the cavities. One possibility is the use of microcavities in photonic crystals [18, 19]. A further option is to use integrated optical circuits, where in principle arbitrary waveguide forms can be created with high precision by laser engraving in the silica substrate [20]. They have been successfully used for the demonstration of a quantum gate operation [21], creation of classical and quantum correlations [22, 23], multi-photon entangled state preparation [24], quantum random walk [25], discrete Fourier transform [26] or Bloch oscillations [27].

While it has been demonstrated that cavities can be fabricated by creating the Bragg grating during the laser writing process [28, 29], there is now an active experimental effort to combine the waveguides with atomic microtraps on a single device [30, 31].

Motivated by these developments and the prospects of studying many-body physics using integrated optical circuits with trapped atoms, we theoretically analyze the non-equilibrium physics of such system, which we take to be a two-dimensional cavity array, where each cavity hosts a number of atoms. We first derive the effective Hamiltonian describing the system in section 2. Using this Hamiltonian we find mean-field (MF) equations of



motion whose solutions feature rich dynamics, including bistable behavior, Ising like phase transition or the occurrence of limit cycles, which we discuss in detail in section 3. We then conclude in section 4.

2. The model

We consider a two-dimensional array of identical three-level atoms each coupled to a horizontal cavity mode a_i and vertical cavity mode b_ν with the same coupling constant g , figure 1(a). From here on we use latin and greek indices to denote horizontal and vertical degrees of freedom respectively. In order to simplify the notation, we omit the sum symbols and summation over occurrences of indices labeling the *spatial* position is understood (i.e. $o_j \equiv \sum_j o_j$). We will use indices with bar, e.g. \bar{j} , when we want to emphasize that the index \bar{j} is fixed and is not summed over. The atomic level structure is depicted in figure 1(b). The $|g\rangle - |e\rangle$ transition is coupled to cavity mode a_i (b_ν) with frequency ω_i (ω_ν). The $|s\rangle - |e\rangle$ transition is coupled using a strong transverse classical light, whose direction of propagation is perpendicular to the plane of the cavity array and which has Rabi frequency Ω , carrier frequency ω_T and detuning Δ_e .

The main reason for choosing a three level Λ system is that it allows, in the limit of large Δ_e , to adiabatically eliminate the excited state $|e\rangle$ in order to avoid the losses due to the spontaneous emission and to obtain an effective Hamiltonian in the ground state manifold subspace. This effective Hamiltonian features a high tunability, see below. We now proceed with the derivation of the effective Hamiltonian. We discuss the issue of cavity gain and loss later in section 3.

The system is described by the Hamiltonian (in the rotating wave approximation)

$$\begin{aligned}
 H = & \omega_i a_i^\dagger a_i + \omega_\nu b_\nu^\dagger b_\nu + \omega_e |e\rangle_{i\nu} \langle e|_{i\nu} + \omega_s |s\rangle_{i\nu} \langle s|_{i\nu} \\
 & + \Omega (|e\rangle_{i\nu} \langle s|_{i\nu} e^{-i\omega_T t} + |s\rangle_{i\nu} \langle e|_{i\nu} e^{i\omega_T t}) \\
 & + g (|e\rangle_{i\nu} \langle g|_{i\nu} (a_i + b_\nu) + \text{h.c.}).
 \end{aligned} \tag{2.1}$$

The excited state $|e\rangle$ can be adiabatically eliminated in the standard way, see appendix A for details. The resulting effective Hamiltonian reads

$$H = \Delta_i a_i^\dagger a_i + \Delta_\nu b_\nu^\dagger b_\nu + \frac{\tilde{\omega}_{a,i\nu}}{2} \sigma_{i\nu}^z + \tilde{g} (\sigma_{i\nu}^+ (a_i + b_\nu) + \text{h.c.}) + F_{i\nu}, \tag{2.2}$$

where σ are the usual Pauli matrices in the $\{|s\rangle, |g\rangle\}$ basis,

$$\begin{aligned}
 \tilde{\omega}_{a,i\nu} &= \Delta_s - \frac{\Omega^2 - g^2 (a_i^\dagger + b_\nu^\dagger)(a_i + b_\nu)}{\Delta_e} \\
 \tilde{g} &= -\frac{g\Omega}{\Delta_e} \\
 F_{i\nu} &= \frac{1}{2} \left(\Delta_s - \frac{\Omega^2 + g^2 (a_i^\dagger + b_\nu^\dagger)(a_i + b_\nu)}{\Delta_e} \right)
 \end{aligned} \tag{2.3}$$

and $\Delta_x = \omega_x - \omega_{\text{aux}}$, $x = \bar{i}, \bar{v}, e$ and $\Delta_s = \omega_s - (\omega_{\text{aux}} - \omega_T)$. Here, ω_{aux} is an auxiliary frequency used in the adiabatic elimination which, in principle, can be chosen arbitrarily. Its physical motivation and interpretation will be discussed momentarily in section 3.1.

3. MF treatment of the nonlinear JC model

The effective Hamiltonian (2.2) is the starting point in this section, where we analyze the effect of cavity loss and pump on the dynamics. Also, it can be seen from the form of the Hamiltonian and the expressions (2.3), that the parameters are highly tunable through varying Ω , Δ_e and Δ_s .

We first perform a MF analysis of a simpler system with only one cavity mode, which already contains rich physics as we show below. We then turn back to the multi-mode two-dimensional setup in section 3.2.

3.1. Single cavity

Cavity without pump

Lets first consider the Hamiltonian (2.2) in the single cavity limit with the single mode a , i.e. we drop the indices i and ν (see figure 1(c)). We rewrite the Hamiltonian as

$$H = \left(\frac{\Delta_{\text{at}}}{2} + \lambda a^\dagger a \right) \Sigma^z + \Delta_{\text{ph}} a^\dagger a + \tilde{g} (\Sigma^+ a + a^\dagger \Sigma^-), \quad (3.1)$$

where $\Sigma_i^{\pm} = \sigma_i^{\pm}$ are the global spin operators and

$$\begin{aligned} \Delta_{\text{at}} &= \Delta_s - \frac{\Omega^2}{\Delta_e} \\ \Delta_{\text{ph}} &= \Delta - \frac{g^2}{2\Delta_e} = \omega - \omega_{\text{aux}} - \frac{g^2}{2\Delta_e} \\ \lambda &= -\frac{g^2}{2\Delta_e}, \end{aligned} \quad (3.2)$$

with ω being the cavity frequency. Note that the model (3.1) without the λ term is the usual JC model [32]. Below we show, that the λ term is indeed at the origin of intriguing system dynamics (see also [33–40] for various other nonlinear extensions of the JC model).

One can now derive the equation of motion for the operator o according to $\dot{o} = -i[o, H]$, where H is given by (3.1). At the same time, any realistic cavity is subject to a decay of the electromagnetic field into the environment. The dissipation process is typically described by means of a master equation (see e.g. [41]), which corresponds to an extra term in the equation of motion for the cavity mode operator, $\dot{a} \propto -\kappa a$, where κ denotes the cavity loss rate (see figure 1(a)). Introducing the expectation values of the operators $\alpha = \langle a \rangle$, $s = \langle \Sigma^- \rangle$, $w = \langle \Sigma^z \rangle$, it is now straightforward to derive the the MF equations of motion, which read

$$i\dot{\alpha} = (\lambda w + \Delta_{\text{ph}} - i\kappa)\alpha + \tilde{g}s, \quad (3.3a)$$

$$i\dot{s} = (\Delta_{\text{at}} + 2\lambda|\alpha|^2)s - \tilde{g}w\alpha, \quad (3.3b)$$

$$i\dot{w} = 2\tilde{g}(s^*\alpha - \alpha^*s). \quad (3.3c)$$

In the derivation we have used the MF decoupling $\langle a^\dagger \Sigma^- \rangle = \alpha^* w$ and $\langle \Sigma^z a \rangle = w\alpha$. We have also neglected the spin decay on the transition $|g\rangle \rightarrow |s\rangle^2$.

First, we wish to find a steady state solution of the equations of motion (3.3). From (3.3a) and (3.3c) we have

$$\alpha = -\frac{\tilde{g}}{\lambda w + \Delta_{\text{ph}} - i\kappa}s \equiv Cs, \quad (3.4)$$

$$s^*\alpha = \alpha^*s. \quad (3.5)$$

Substituting the equation (3.4) to (3.5) yields

$$C|s|^2 = C^*|s|^2. \quad (3.6)$$

When C is complex, which occurs only for non-zero cavity decay $\kappa \neq 0$, the condition equation (3.6) can be satisfied if $s = 0$ (and consequently $\alpha = 0$), which results in a trivial solution with empty cavity and all spins

² This is a justified assumption in implementations with real atoms as the levels $|g\rangle$, $|s\rangle$ typically belong to some ground state manifold, where only magnetic dipole transitions are allowed between the states of that manifold. In turn, the spin decay is negligible compared to the cavity decay.

down (up) in the steady state. This indicates that in order to obtain some non-trivial physics in the steady state limit and to counteract the cavity losses, one needs to input energy into the system³. In our case a natural choice is either through the pumping of the cavity mode or driving directly the $|g\rangle \rightarrow |s\rangle$ transition. We focus here on the former case only.

Cavity with pump

The cavity pump can be described by the Hamiltonian

$$H_{\text{pump}} = \eta a^\dagger e^{-i\omega_p t} + \eta^* a e^{i\omega_p t}. \quad (3.7)$$

When adding H_{pump} to the system Hamiltonian (2.1), the explicit time dependence in the total Hamiltonian can be removed if the auxiliary frequency ω_{aux} in (2.3) is chosen such that it is the frequency of the cavity pump, $\omega_{\text{aux}} = \omega_p$. The MF equations of motion (3.3) then become

$$\begin{aligned} \dot{\alpha}_R &= -\kappa\alpha_R + (\lambda w + \Delta_{\text{ph}})\alpha_I - \frac{\tilde{g}}{2}s_y + \eta_I \\ \dot{\alpha}_I &= -(\lambda w + \Delta_{\text{ph}})\alpha_R - \kappa\alpha_I - \frac{\tilde{g}}{2}s_x - \eta_R \\ \dot{s}_x &= -2\tilde{g}w\alpha_I - (\Delta_{\text{at}} + 2\lambda|\alpha|^2)s_y \\ \dot{s}_y &= -2\tilde{g}w\alpha_R + (\Delta_{\text{at}} + 2\lambda|\alpha|^2)s_x \\ \dot{w} &= 2\tilde{g}(\alpha_R s_y + \alpha_I s_x), \end{aligned} \quad (3.8)$$

where we have introduced real variables through $\alpha = \alpha_R + i\alpha_I$, $s = 1/2(s_x - is_y)$ and $\eta = \eta_R + i\eta_I$.

Note that the equations (3.8) imply the conservation of the total spin

$$w^2 + 4|s|^2 = N^2, \quad (3.9)$$

where N is the total number of the spins. This can be easily verified as

$$\partial_t S^2 = \partial_t (s_x^2 + s_y^2 + w^2) = \partial_t (4ss^* + w^2) = 0, \quad (3.10)$$

where we have parametrized the total spin as $\vec{S} = (s_x, s_y, w)$.

The set of equations (3.8) will serve as the starting point for most of the analysis in this section. Ultimately, we are interested in the effect of the nonlinear term proportional to λ , which makes the system described by (3.1) qualitatively different from the usual JC model. Due to the complexity of the problem, we start by studying a simpler situation in section 3.1.1 by putting $\lambda = 0$ (and as we will see also $\Delta_{\text{at}} = 0$). Building upon the $\lambda = 0$ solution, we then analyze the $\lambda \neq 0$ and $\Delta_{\text{at}} \neq 0$ situations in sections 3.1.2 and 3.1.3 respectively.

3.1.1. $\lambda = 0$ regime

In order to investigate the steady state solutions, we first put $\lambda = 0$ to further simplify the problem (see e.g. [8, 42, 43] for related studies of the phase diagram of the JC and Dicke models). We then turn back to the situation with $\lambda \neq 0$ in the next section. With these simplifications, the real equations for the steady state read

$$0 = -\kappa\alpha_R + \Delta_{\text{ph}}\alpha_I - \frac{\tilde{g}}{2}s_y + \eta_I, \quad (3.11a)$$

$$0 = -\Delta_{\text{ph}}\alpha_R - \kappa\alpha_I - \frac{\tilde{g}}{2}s_x - \eta_R, \quad (3.11b)$$

$$0 = -2\tilde{g}w\alpha_I - \Delta_{\text{at}}s_y, \quad (3.11c)$$

$$0 = -2\tilde{g}w\alpha_R + \Delta_{\text{at}}s_x \quad (3.11d)$$

$$0 = 2\tilde{g}(\alpha_R s_y + \alpha_I s_x). \quad (3.11e)$$

Solving the set (3.11a) and (3.11b) for α_R , α_I and substituting to (3.11c)–(3.11e) we obtain for the spin conservation (3.9)

$$w^2 - N^2 = \frac{-4\tilde{g}^2 w^2 \eta^2}{g^4 w^2 + \Delta_{\text{at}}^2 \kappa^2 + \Delta_{\text{ph}} \Delta_{\text{at}} (2g^2 w + \Delta_{\text{ph}} \Delta_{\text{at}})}, \quad (3.12)$$

where $\eta^2 = \eta_R^2 + \eta_I^2$. This is a 4th order polynomial for w and its solutions in terms of radicals can in principle be found yielding rather complicated expressions, which are not of much practical use. Instead, we will analyze the properties of (3.12) as follows. Since $w/N \in [-1, 1]$, the lhs of (3.12) is non-positive, namely $(w/N)^2 - 1 \in [-1, 0]$. At the same time the nominator of the rhs is clearly non-positive, so that the non-

³ Note that this is in sharp contrast with the full Dicke Hamiltonian, where the presence of the counterrotating terms guarantees non-trivial solutions even in the absence of the pumping [4, 8]. In the model we study, the absence of the counterrotating terms is a direct consequence of applying the rotating wave approximation, as the cavity modes and the $|g\rangle \rightarrow |e\rangle$ transition are taken to be at optical frequencies.

positivity of the rhs requires the quadratic polynomial in the denominator to be non-negative. We thus have

$$g^4 w^2 + \Delta_{\text{at}}^2 \kappa^2 + \Delta_{\text{ph}} \Delta_{\text{at}} (2g^2 w + \Delta_{\text{ph}} \Delta_{\text{at}}) \geq 0. \quad (3.13)$$

The roots of this polynomial read

$$w_{\pm} = \frac{-2\tilde{g}^2 \Delta_{\text{ph}} \Delta_{\text{at}} \pm \sqrt{-4\tilde{g}^4 \Delta_{\text{at}}^2 \kappa^2}}{2\tilde{g}^4}. \quad (3.14)$$

It can be immediately seen that for $\kappa \neq 0$, any $\Delta_{\text{at}} \neq 0$ would yield imaginary w . This however does not mean that a solution of the original constraint (3.12) is not possible for both $\kappa, \Delta_{\text{at}} \neq 0$. Rather, it means that the point $\Delta_{\text{at}} = 0$ has some particular properties which we will investigate further.

For $\Delta_{\text{at}} = 0$, the steady state equations read

$$\kappa \alpha_{\text{R}} - \Delta_{\text{ph}} \alpha_{\text{I}} = -\frac{\tilde{g}}{2} s_y + \eta_{\text{I}}, \quad (3.15a)$$

$$\Delta_{\text{ph}} \alpha_{\text{R}} + \kappa \alpha_{\text{I}} = -\frac{\tilde{g}}{2} s_x - \eta_{\text{R}}, \quad (3.15b)$$

$$0 = w \alpha_{\text{I}}, \quad (3.15c)$$

$$0 = w \alpha_{\text{R}}, \quad (3.15d)$$

$$\alpha_{\text{R}} s_y = -\alpha_{\text{I}} s_x. \quad (3.15e)$$

Equations (3.15c) and (3.15d) imply either $w = 0$ or $\alpha_{\text{R}} = \alpha_{\text{I}} = 0$. If $\alpha_{\text{R}} = \alpha_{\text{I}} = 0$, the equations can be readily solved to yield

$$\begin{aligned} s_x &= -\frac{2\eta_{\text{R}}}{\tilde{g}} \\ s_y &= \frac{2\eta_{\text{I}}}{\tilde{g}} \\ w &= \pm \sqrt{1 - \frac{4\eta^2}{\tilde{g}^2}}. \end{aligned} \quad (3.16)$$

These solutions are valid only for $\tilde{g} \geq 2\eta/N \equiv g_1^*$.

In the case where $w = 0$, we can use the spin conservation (3.9) to parametrize the spin as $s_x = N \cos \theta$, $s_y = -N \sin \theta$. (3.15e) then becomes

$$\frac{\alpha_{\text{I}}}{\alpha_{\text{R}}} = \tan \theta. \quad (3.17)$$

Next we can express $\alpha_{\text{R}}, \alpha_{\text{I}}$ from (3.15a) and (3.15b) as

$$\begin{aligned} \alpha_{\text{R}} &= \frac{1}{2(\Delta_{\text{ph}}^2 + \kappa^2)} [\kappa(\tilde{g}N \sin \theta + 2\eta_{\text{I}}) - \Delta_{\text{ph}}(\tilde{g}N \cos \theta + 2\eta_{\text{R}})], \\ \alpha_{\text{I}} &= -\frac{1}{2(\Delta_{\text{ph}}^2 + \kappa^2)} [\Delta_{\text{ph}}(\tilde{g}N \sin \theta + 2\eta_{\text{I}}) + \kappa(\tilde{g}N \cos \theta + 2\eta_{\text{R}})]. \end{aligned} \quad (3.18)$$

Substituting these expressions to (3.17) yields the condition for the angle θ which determines the solution and can be found numerically. In order to proceed further analytically, we put $\eta_{\text{I}} = 0$. The equation (3.17) can then be cast in the form

$$z + \kappa \cos \theta = \Delta_{\text{ph}} \sin \theta \quad (3.19)$$

or equivalently

$$(\kappa^2 + \Delta_{\text{ph}}^2) \cos^2 \theta + 2\kappa z \cos \theta + z^2 - \Delta_{\text{ph}}^2 = 0 \quad (3.20)$$

with solutions

$$\cos \theta_{\pm} = \frac{-2\kappa z \pm \sqrt{4\kappa^2 z^2 - 4(z^2 - \Delta_{\text{ph}}^2)(\kappa^2 + \Delta_{\text{ph}}^2)}}{2(\kappa^2 + \Delta_{\text{ph}}^2)}, \quad (3.21)$$

where $z = \kappa \tilde{g} N / (2\eta_{\text{R}})$. Assuming $\Delta_{\text{ph}} > 0$, we plot the solutions (3.21) as a function of \tilde{g} in figure 2. When tuning \tilde{g} , the non-negativity of the discriminant in (3.21) determines the maximal coupling g_2^* up to which the real solutions (3.21) can be found. It reads

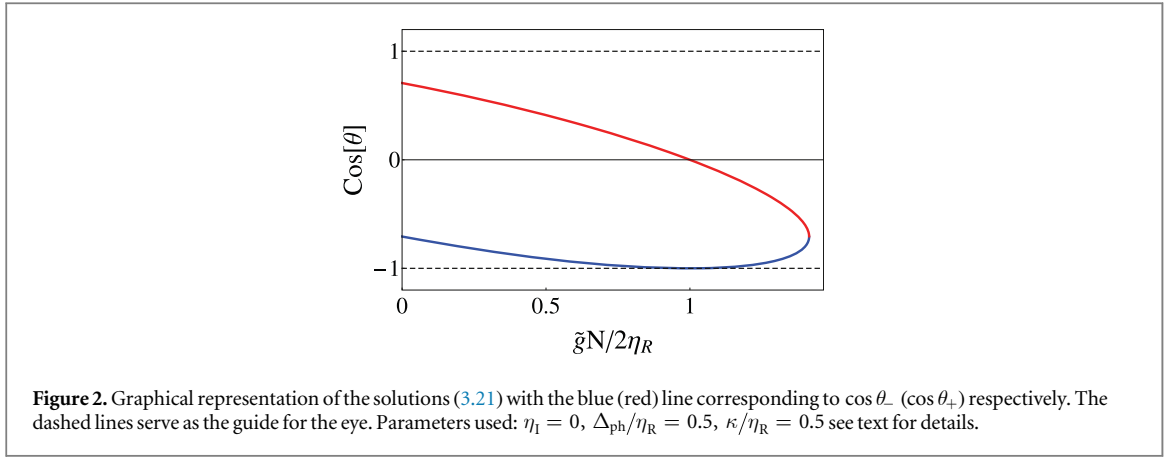


Figure 2. Graphical representation of the solutions (3.21) with the blue (red) line corresponding to $\cos \theta_-$ ($\cos \theta_+$) respectively. The dashed lines serve as the guide for the eye. Parameters used: $\eta_I = 0$, $\Delta_{ph}/\eta_R = 0.5$, $\kappa/\eta_R = 0.5$ see text for details.

$$g_2^* = \frac{2\eta_R}{N} \sqrt{1 + \left(\frac{\Delta_{ph}}{\kappa}\right)^2} \quad (3.22)$$

and corresponds to the value of \tilde{g} , where $\cos \theta_- = \cos \theta_+$, which can be seen as the rightmost point in figure 2.

Note, that the region $\tilde{g} \in [g_1^*, g_2^*]$ admits both solutions (3.16) and ((3.18) and (3.21)) indicating a bistable behavior. In what follows we refer to the points g_1^* , g_2^* as transition points.

Stability study

The stability analysis of the steady state solutions (3.16) and ((3.18) and (3.21)) is performed in a standard way by linearizing (3.8) around the solutions, i.e. expressing the variables as $v = \bar{v} + \delta v$, where \bar{v} denotes the steady state solution. Formally this yields the linearized equations of motion

$$\dot{\delta v} = M\delta v + b, \quad (3.23)$$

where $\delta v = (\delta\alpha_R, \delta\alpha_I, \delta s_x, \delta s_y, \delta w)^T$. Instability of the solutions is indicated by the positivity of the real part of the maximum eigenvalue of M , see appendix B for details. It is noteworthy, that in any steady state solution we have either $\bar{w} = 0$ or $\bar{\alpha}_R = \bar{\alpha}_I = 0$ and consequently the characteristic polynomial of the matrix M becomes

$$|M - y\mathbb{I}| = yp(y^4), \quad (3.24)$$

where $p(y^4)$ is some polynomial which is 4th order in y . We thus always have one eigenvalue $y = 0$, which is simply the consequence of the spin conservation law (3.9). Examining numerically the negativity of real part of the roots of $p(y^4)$ we identify stable and unstable solutions. These are depicted by solid and dashed lines respectively in the steady state phase diagram figure 3, where we plot $|\alpha|^2$ and w as functions of the coupling \tilde{g} . The region II exhibits bistable behavior, whereas regions I and III admit only a single stable solution. We have also verified, by numerically solving the dynamical equations (3.8), that they indeed evolve into the steady state solutions (3.16), (3.18) and (3.21) (not shown).

A comment on the validity of the MF approach is in order. At first sight, the presence of bistable regions in the MF solutions seems to be incompatible with the expected existence of a unique stationary state as it suggests that there are two different states for a given value of the coupling \tilde{g} . However, such situation commonly occurs in the MF treatment of many-body systems that feature a first order phase transition. It can be encountered, for example, in the context of the classical van der Waals gas, described by a MF equation of state [44] or, more recently, in the context of Rydberg gases, where optical bistability was observed in qualitative agreement with the MF predictions [45].

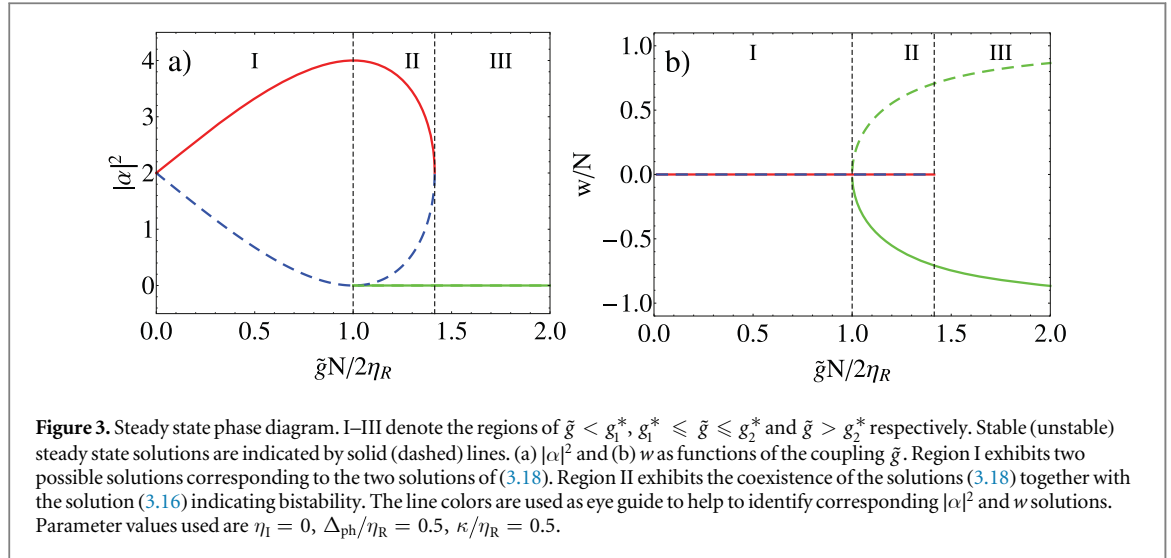
In the next two sections we investigate how the inclusion of the non-zero atomic detuning Δ_{at} and non-zero coupling λ terms modifies the $\Delta_{at} = \lambda = 0$ solution.

3.1.2. $\lambda \neq 0$ regime

In order to shed light on the effect of the λ term independently of the Δ_{at} term, we keep $\Delta_{at} = 0$. From (3.8) the steady state equations read

$$-\kappa\alpha_R + (\Delta_{ph} + \lambda w)\alpha_I = -\frac{\tilde{g}}{2}s_y + \eta_I, \quad (3.25a)$$

$$(\Delta_{ph} + \lambda w)\alpha_R + \kappa\alpha_I = -\frac{\tilde{g}}{2}s_x - \eta_R, \quad (3.25b)$$



$$\lambda|\alpha|^2 s_y = -\tilde{g}w\alpha_I, \quad (3.25c)$$

$$\lambda|\alpha|^2 s_x = \tilde{g}w\alpha_R, \quad (3.25d)$$

$$\alpha_R s_y = -\alpha_I s_x. \quad (3.25e)$$

Finding the steady state solution encompasses solving a 6th order polynomial equation, which in general can only be done numerically. Nevertheless, some information about the steady state solution can be obtained analytically as we now describe.

Substituting s_x , s_y from (3.25c) and (3.25d) to the spin conservation (3.9), we get the following condition for w

$$w^2 = \frac{\lambda^2 |\alpha|^2}{\lambda^2 |\alpha|^2 + \tilde{g}^2} N^2. \quad (3.26)$$

Since $\lambda^2 |\alpha|^2 / (\lambda^2 |\alpha|^2 + \tilde{g}^2) \leq 1$, the relation (3.26) indicates, that there is no further instability ($w^2 > N^2$) when changing λ , i.e. there are also two transition points g_1^* , g_2^* for $\lambda \neq 0$.

Next, one can find solutions for asymptotic values of the parameter λ . Clearly, for $\lambda \rightarrow 0$, one should recover the solutions of section 3.1.1. On the other hand, for $\lambda \rightarrow \infty$, one can look for solution by substituting a perturbative expansion for all the variables of the form $v = \sum_{n=0}^{\infty} \lambda^{-n} v^{(n)}$. We provide the details of this expansion in appendix C. In order to simplify the analytic expressions we take the imaginary part of the pump to be zero, $\eta_I = 0$. In this case, to leading order in λ , the solutions read $\alpha_R^{(0)} = \alpha_I^{(0)} = s_y^{(0)} = 0$ and

$$s_x^{(0)} = -\frac{2\eta}{\tilde{g}} \quad \text{for } \tilde{g} \geq \frac{2\eta}{N}, \quad (3.27a)$$

$$s_x^{(0)} = \frac{\eta}{\tilde{g}} \left(1 - \sqrt{1 + \frac{2\tilde{g}^2 N^2}{\eta^2}} \right) \quad \text{for } \tilde{g} \leq \frac{2\eta}{N}. \quad (3.27b)$$

Interestingly, since the spin conservation implies $|s_x^{(0)}| \leq N$, the two solutions (3.27a) and (3.27b) yield the same transition point $g_1^* = 2\eta/N$. This is in contrast to the $\lambda = 0$ case, where two distinct transition points g_1^* , g_2^* were identified. Here, the solution (3.27a) is valid for $\tilde{g} \geq 2\eta/N$, whereas (3.27b) is valid for $\tilde{g} \leq 2\eta/N$. This indicates that the transition point g_2^* approaches g_1^* as λ is increased until they become identical in the $\lambda \rightarrow \infty$ limit. In other words, the region II of the phase diagram figure 3 is shrinking to zero width as λ is increased. For the intermediate values of λ , the values of g_1^* , g_2^* can be found numerically (selecting only the physically meaningful solutions of the 6th order polynomial, corresponding to the real values of w). We plot the shrinking of the $[g_1^*, g_2^*]$ region in figure 4(a).

In figure 5 we show the phase diagrams for $|\alpha|^2$ and w for increasing values of λ . It can be seen from figures 5(a) and (b), that for some values of λ , there is a region with no apparent stable solution. In nonlinear systems, this is typically a signature of the appearance of limit cycles which occur through a Hopf bifurcation as the system leaves the stable fixed point by changing the coupling \tilde{g} [46]. We represent a limit cycle corresponding to the point A in figure 5(b) as the time evolution of the global spin in figure 6.

One can carry the analysis further in the large λ limit and look for asymptotic solutions in the vicinity of the unique transition point g_1^* . The first non-trivial contribution to α is of order λ^{-1} and given by (see appendix C)

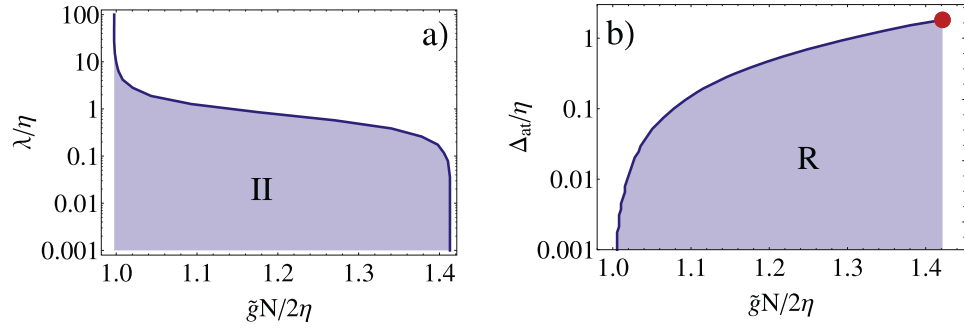


Figure 4. (a) $\lambda \neq 0$ regime: the shaded area represents the shrinking of the $[g_1^*, g_2^*]$ region (denoted as region II, see also figure 3) with increasing λ . (b) $\lambda = 0$ regime: the shaded area corresponds to the region of \tilde{g} exhibiting four solutions for a given \tilde{g} (denoted as region R, see also figures 7(a) and (b)) as Δ_{at} is increased. The critical value of Δ_{at} at which the four simultaneous solutions cease to exist is represented by the red circle.

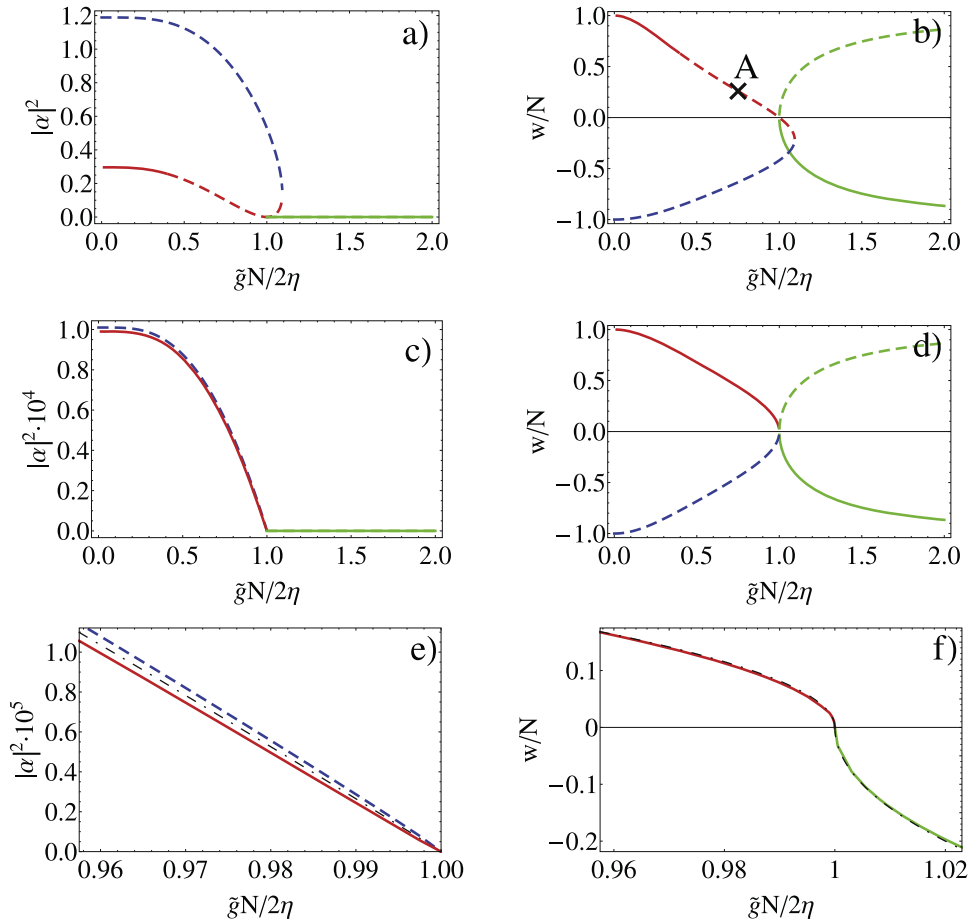
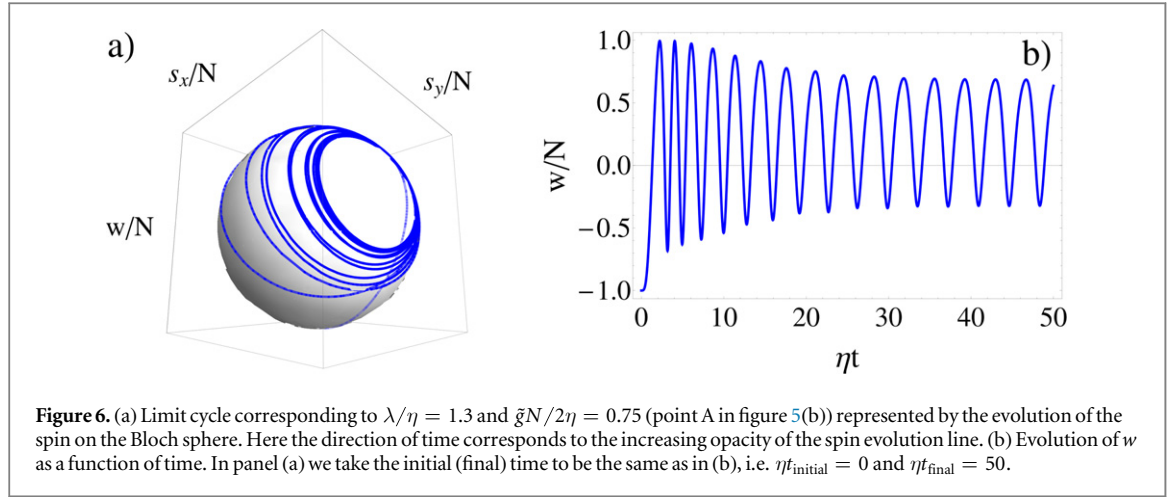


Figure 5. The cavity field $|\alpha|^2$ and the expectation value w of the spin as functions of the coupling strength \tilde{g} for various values of the nonlinearity λ : $\lambda/\eta = 1.3$ in (a) and (b) and $\lambda/\eta = 100$ in (c) and (d). Stable (unstable) solutions are shown as solid (dashed) lines. Panels (e) and (f) show a magnification of the data shown in (c) and (d) in the vicinity of the critical point \tilde{g}_1^* . The dashed dotted lines represent the analytic scaling solutions (3.28) and (3.29) respectively. In all plots the line colors are used as eye guide to help to identify the corresponding $|\alpha|^2$ and w solutions. In panel (b), the point A denotes the solution corresponding to the limit cycle, see text and figure 6 for more details.

$$|\alpha|^2 = \frac{1}{\lambda^2} \frac{1}{27} (g_1^* - \tilde{g}) \left(\frac{10\eta}{N} + 13\tilde{g} \right) + O((g_1^* - \tilde{g})^{\frac{3}{2}}), \quad (3.28)$$

which is valid for $\tilde{g} \leq g_1^*$ as the solution for $\tilde{g} > g_1^*$ is trivial ($\alpha = 0$). The scaling of the spin observables, say w , is simply obtained from the expansion of (3.27), which read



$$w^{(0)} = \pm N \sqrt{\frac{N}{\eta}} \sqrt{\tilde{g} - g_1^*} + O\left((\tilde{g} - g_1^*)^{\frac{3}{2}}\right) \text{ for } \tilde{g} \geq g_1^*, \quad (3.29a)$$

$$w^{(0)} = \pm N \sqrt{\frac{N}{3\eta}} \sqrt{g_1^* - \tilde{g}} + O\left((g_1^* - \tilde{g})^{\frac{3}{2}}\right) \text{ for } \tilde{g} \leq g_1^*, \quad (3.29b)$$

where the sign \pm depends on the branch of w considered. Note that the scaling $(\tilde{g} - g_1^*)^{1/2}$ is characteristic of the Ising type model with \mathbb{Z}_2 symmetry and g_1^* is the phase transition critical point. We compare the asymptotic solutions (3.28) and (3.29) with the solutions obtained by numerically solving (3.25) in figures 5(e) and (f).

3.1.3. $\lambda = 0$ regime with $\Delta_{\text{at}} \neq 0$

Our next aim is to explore the effect of the Δ_{at} term on the solution in the absence of the nonlinear λ term. In this case, the structure of the solutions is dictated by the 4th order w polynomial (3.12) and it is straightforward to obtain the solutions numerically. For small Δ_{at} , there is a range of \tilde{g} -values, which admits four solutions (corresponding to four real solutions of (3.12)), figures 7(a) and (b). As Δ_{at} is increased, this region eventually disappears leaving us with only two solutions for all values of \tilde{g} . The latter limit can be simply understood from the MF equations (3.11). In the large Δ_{at} limit, the leading contribution comes from (3.11c) and (3.11d) with the trivial solutions $s_x = s_y = 0$ implying $w/N = \pm 1$. The evolution of the solutions towards this large Δ_{at} limit can be seen rather clearly from figure 7(d). Next, we have determined numerically the size of the region of \tilde{g} admitting four solutions as a function of Δ_{at} . This is shown in figure 4(b). Clearly, there is some maximal Δ_{at} after which there are only two possible solutions, as discussed. This limiting value is represented by the red circle in figure 4(b). Note, that the inclusion of the Δ_{at} term also lifts the transition point g_1^* , i.e. that all solutions are smooth in the vicinity of g_1^* .

3.2. Multiple cavities

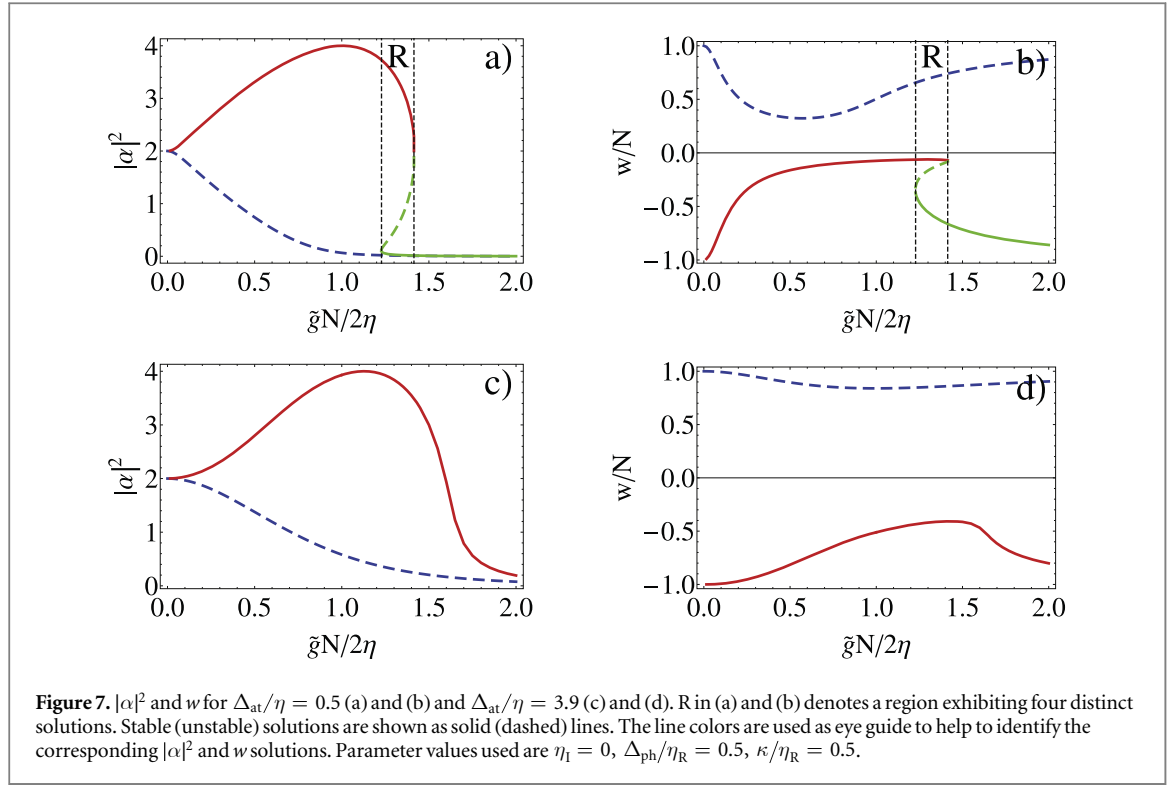
In this section we seek the generalization of the single cavity case to higher dimensional geometries. For concreteness, we consider a 2D square geometry depicted schematically in figure 1(a).

Before diving in the details of the analysis, we motivate this section by asking whether a 2D square geometry offers MF solutions which are qualitatively different from the 1D case studied above. For example, it was shown in [47] in the context of laser driven and interacting Rydberg gases on a square lattice, that a homogeneous system admits a MF solution that breaks the lattice symmetry and exhibits antiferromagnetic (AF) order.

We start our discussion by deriving the 2D MF equations, which follow from the operator equations of motion (A.10) given by the Hamiltonian (2.2). Repeating the argument yielding (3.6) leads to the same conclusion, namely that in the absence of the cavity pump the system possesses only a trivial solution, where the cavity modes are empty and all the spins are down (up). We first focus on the situation without the nonlinear term, $\lambda = 0$. In the following, we consider the cavity pump η , the cavity decay κ and the spin decay γ to be the same for all cavity modes and all spins respectively (the motivation for adding the spin decay will become clear shortly). We obtain the set of MF equations

$$i\dot{\alpha}_{\bar{i}} = (\Delta_{\text{ph}}^{\bar{i}} - i\kappa + \lambda w_{i\nu}^{\bar{i}})\alpha_{\bar{i}} + \tilde{g}_a s_{i\nu}^{\bar{i}} + \eta + \lambda\beta_{\nu}(w_{i\nu}^{\bar{i}} - \mathbb{1}), \quad (3.30a)$$

$$i\dot{\beta}_{\bar{\nu}} = (\Delta_{\text{ph}}^{\bar{\nu}} - i\kappa + \lambda w_{i\nu}^{\bar{\nu}})\beta_{\bar{\nu}} + \tilde{g}_b s_{i\nu}^{\bar{\nu}} + \eta + \lambda\alpha_i(w_{i\nu}^{\bar{\nu}} - \mathbb{1}), \quad (3.30b)$$



$$i\dot{s}_{\bar{i}\bar{\nu}} = \left(\Delta_{\text{at}} - i\frac{\gamma}{2} + 2\lambda(\alpha_{\bar{i}}^* + \beta_{\bar{\nu}}^*)(\alpha_{\bar{i}} + \beta_{\bar{\nu}}) \right) s_{\bar{i}\bar{\nu}} - (\tilde{g}_a \alpha_{\bar{i}} + \tilde{g}_b \beta_{\bar{\nu}}) w_{\bar{i}\bar{\nu}}, \quad (3.30c)$$

$$i\dot{w}_{\bar{i}\bar{\nu}} = 2[s_{\bar{i}\bar{\nu}}^*(\tilde{g}_a \alpha_{\bar{i}} + \tilde{g}_b \beta_{\bar{\nu}}) - \text{h.c.}] - i\gamma(w_{\bar{i}\bar{\nu}} + 1). \quad (3.30d)$$

Here, we allowed for the couplings \tilde{g} and the photon detunings Δ_{ph} to be different for rows and columns while taking the remaining parameters to be the same for all cavities.

Clearly, the 2D system provides higher tunability by enlarging the parameter space. In the following, we study the problem considering two different perspectives, namely a homogeneous and cluster-MF ansatz.

3.2.1. Square array with homogeneous MF

In order to further simplify the MF equations (3.30), it is reasonable to use the typical MF ansatz, namely that, due to translational symmetry in either direction of the 2D array (along rows or columns), the corresponding cavity fields are the same ($\alpha_{\bar{i}} = \alpha$, $\beta_{\bar{\nu}} = \beta$ for all \bar{i} , $\bar{\nu}$) as well as the spins, $s_{\bar{i}\bar{\nu}} = s$, $w_{\bar{i}\bar{\nu}} = w$. We can now proceed along similar lines as in section 3.1.1 in order to characterize the solutions. In analogy to section 3.1.1 we start with $\Delta_{\text{at}} = \lambda = 0$ and we neglect the spin decay for the moment, $\gamma = 0$. Considering $\Delta_{\text{ph}}^{\bar{i}} \equiv \Delta_{\text{ph}}^a$ and $\Delta_{\text{ph}}^{\bar{\nu}} \equiv \Delta_{\text{ph}}^b$ for all \bar{i} , $\bar{\nu}$ to be the same along rows or columns, the steady state MF equations (3.30) simplify to

$$0 = (\Delta_{\text{ph}}^a - i\kappa)\alpha + \tilde{g}_a N_C s + \eta, \quad (3.31a)$$

$$0 = (\Delta_{\text{ph}}^b - i\kappa)\beta + \tilde{g}_b N_R s + \eta, \quad (3.31b)$$

$$0 = (\tilde{g}_a \alpha + \tilde{g}_b \beta)w, \quad (3.31c)$$

$$0 = s^*(\tilde{g}_a \alpha + \tilde{g}_b \beta) - \text{h.c.} \quad (3.31d)$$

Here N_R (N_C) is the number of rows (columns) respectively. Since the fields β and α share the same spin s , β can be given directly in terms of the field α . The situation is then essentially equivalent to the single cavity case analyzed in section 3.1, though with some extra tunability provided by larger number of parameters. For example, one can find the transition point \tilde{g}_1^* by combining (3.31a) and (3.31b) with $\tilde{g}_a \alpha + \tilde{g}_b \beta = 0$, which is one possible solution of (3.31c). The transition point then corresponds to the situation where $w = 0$, i.e.

$$|s|^2 = \frac{1}{4} = |\eta|^2 \frac{(\Delta_{\text{ph}}^b \tilde{g}_a + \Delta_{\text{ph}}^a \tilde{g}_b)^2 + (\tilde{g}_a + \tilde{g}_b)^2 \kappa^2}{(\Delta_{\text{ph}}^b \tilde{g}_a^2 N_C + \Delta_{\text{ph}}^a \tilde{g}_b^2 N_R)^2 + (\tilde{g}_a^2 N_C + \tilde{g}_b^2 N_R)^2 \kappa^2}. \quad (3.32)$$

One can then identify the critical value of e.g. \tilde{g}_a for all other parameters fixed. As a consistency check, it is easy to verify that one recovers the single cavity expression $\tilde{g}_1^* = 2|\eta|/N$ by omitting all the 'b' variables and setting $N_R = 1$ and $N_C = N$, the number of spins.

3.2.2. Square array with cluster-MF

In the previous section, we have used the homogeneous MF ansatz and found that the solutions correspond effectively to a single cavity case with no intriguing spin configurations. However, it is known and was shown e.g. in [47], that a suitable MF ansatz can lead to a non-trivial configuration (such as antiferromagnetic ordering) even if the steady state MF equations are completely symmetric under exchange of the spin variables. Here we will consider the simplest possible case of such ansatz where a cluster is formed by two adjacent inequivalent spins s_1, w_1 and s_2, w_2 . On the other hand, since the fields α and β couple to both s_1 and s_2 in the same way, we take α, β to be the same along rows (columns) due to translational symmetry. In order to simplify the equations, we now take all the parameters to be the same along all rows and columns (i.e. we set $\tilde{g}_a = \tilde{g}_b = \tilde{g}$ and $\Delta_{\text{ph}}^{\tilde{i}} = \Delta_{\text{ph}}^{\tilde{v}} = \Delta_{\text{ph}}$ for all \tilde{i}, \tilde{v}). Starting with the simplest case $\Delta_{\text{at}} = \lambda = 0$, the MF equations (3.30) become

$$0 = (\Delta_{\text{ph}} - i\kappa)\alpha + \frac{\tilde{g}\sqrt{N}}{2}(s_1 + s_2) + \eta, \quad (3.33a)$$

$$0 = (\Delta_{\text{ph}} - i\kappa)\beta + \frac{\tilde{g}\sqrt{N}}{2}(s_1 + s_2) + \eta, \quad (3.33b)$$

$$0 = i\frac{\gamma}{2}s_{\tilde{j}} + \tilde{g}(\alpha + \beta)w_{\tilde{j}}, \quad (3.33c)$$

$$0 = 2\tilde{g}[s_{\tilde{j}}^*(\alpha + \beta) - \text{h.c.}] - i\gamma(w_{\tilde{j}} + 1), \quad (3.33d)$$

where $\tilde{j} = 1, 2$ labels the different spins of the cluster and we assumed that each field α, β couples to the same number of spins 1 and 2, hence the factor $1/2$ in the second term of (3.33a) and (3.33b). Here N denotes the total number of spins, i.e. there are \sqrt{N} spins along rows and along columns.

Until now we did not comment on the spin conservation in the 2D case. Going back to the most general situation, where every individual spin and cavity mode is described in terms of the corresponding MF variables $s_{\tilde{i}\tilde{v}}, w_{\tilde{i}\tilde{v}}$ and $\alpha_{\tilde{i}}, \beta_{\tilde{v}}$ respectively, the system evolves according to the full set of the 2D MF equations (3.30). One can verify by means of (3.10) that (3.30) actually imply both local

$$w_{\tilde{i}\tilde{v}}^2 + 4|s_{\tilde{i}\tilde{v}}|^2 = 1, \quad \forall \tilde{i}, \tilde{v} \quad (3.34)$$

and global spin conservation

$$W^2 + 4|\Sigma|^2 = N^2, \quad (3.35)$$

provided $\gamma = 0$. In (3.35) $W = w_{i\bar{i}}$ and $\Sigma = s_{i\bar{i}}$ are the global spin components. One should appreciate that (3.35) actually prevents any spin configuration incompatible with it, including the AF order (which corresponds to $w_1 = -w_2$ when using the here considered cluster-MF). In order to see e.g. the AF order, such as in [47], one needs to break the global spin conservation. This is achieved by the inclusion of the spin decay γ , which breaks both the local and global spin conservations (3.34) and (3.35).

Allowing for the spin decay, we can now ask, whether the solution of the cluster-MF equation (3.33) features any non-trivial spin configuration ($w_1 \neq w_2$). First, it follows from (3.33a) and (3.33b) that $\alpha = \beta$. Next, expressing for $s_{\tilde{j}}$ from (3.33c) and substituting to (3.33d), it can be shown that

$$\frac{w_1 + 1}{w_1} = \frac{w_2 + 1}{w_2} \Rightarrow w_1 = w_2 \quad (3.36)$$

and $s_1 = s_2$ from (3.33c). Once again we find that the structure of the equations (3.33) reduces the problem effectively to the single cavity situation described by a simple set of variables α, s, w and this even when allowing for inequivalent spin configurations and the spin decay.

One might argue, that the equivalence to the single cavity case conjectured in sections 3.2.1 and 3.2.2 is an artefact of taking $\Delta_{\text{at}} = \lambda = 0$. Indeed, when comparing (3.3) and (3.30), one can note that there is a *qualitative* difference between the 1D and 2D situation. Specifically, there is an extra coupling between the a and b modes, the last term in (3.30a) and (3.30b). We have numerically verified that, in a general situation with variables $\alpha, \beta, s_{\tilde{j}}, w_{\tilde{j}}, \tilde{j} = 1, 2$ and $\tilde{g}_a \neq \tilde{g}_b, \Delta_{\text{ph}}^a \neq \Delta_{\text{ph}}^b, \Delta_{\text{at}} \neq 0 \neq \lambda$, the solutions always yield $s_1 = s_2, w_1 = w_2$. For completeness we have included in our numerical analysis also the special cases, where any possible combination of the following conditions can occur: $\tilde{g}_a = \tilde{g}_b, \Delta_{\text{ph}}^a = \Delta_{\text{ph}}^b, \Delta_{\text{at}} = 0, \lambda = 0$.

In summary, the MF equations (3.30) with the ansatz considered in sections 3.2.1 and 3.2.2 on a square lattice reduce to effectively one-dimensional description with no intriguing spin configurations. It would be desirable to perform a beyond MF study of the nonlinear two-dimensional model in order to assess the true nature of the steady state and the corresponding spin and field configurations, including their mutual correlations. Also we did not fully exploit the possibilities offered by the proposed implementation, such as taking different geometries of the array or allowing for disordered coupling strengths, which we leave for further investigations.

4. Conclusion

Motivated by the progress in integrated optical circuits, we have proposed a possible realization of a two-dimensional cavity array with trapped atoms, which is a promising scalable quantum architecture. We derived an effective description of the system in terms of JC like Hamiltonian with highly tunable parameters and extra nonlinear terms. We then analyzed the dynamics of the system using a MF approach. We have found a rich behavior including bistable regions, Ising like phase transition or occurrence of limit cycles through Hopf bifurcations. In the present setup, we have not found conceptual differences between the one and two-dimensional cases at the level of the MF description and with the geometry considered. We hope that the present work lays down grounds for future studies of the cavity arrays realized with integrated optical circuits. The problems which might be addressed in the future are e.g. going beyond MF description, accounting for more exotic geometries or studying effective spin physics as a low energy limit of the presented cavity Hamiltonian.

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Appendix A. Adiabatic elimination

Here we derive the effective Hamiltonian resulting from the adiabatic elimination of the state $|e\rangle$ and starting with the full system Hamiltonian given in (2.1), which we rewrite in the hard-core boson representation with annihilation operators $b_{\mu,\bar{i}\bar{\nu}}$, $\mu = e, s, g$ for the atom at position $\bar{i}\bar{\nu}$

$$\begin{aligned} H = & \omega_i a_i^\dagger a_i + \omega_\nu b_\nu^\dagger b_\nu + \omega_e b_{e,i\nu}^\dagger b_{e,i\nu} + \omega_s b_{s,i\nu}^\dagger b_{s,i\nu} \\ & + \Omega(b_{e,i\nu}^\dagger b_{s,i\nu} e^{-i\omega_T t} + b_{s,i\nu}^\dagger b_{e,i\nu} e^{i\omega_T t}) \\ & + g(b_{e,i\nu}^\dagger b_{g,i\nu}(a_i + b_\nu) + \text{h.c.}). \end{aligned} \quad (\text{A.1})$$

First, we determine the transformation to the rotating frame. We consider a general unitary transformation

$$U = e^{-it(\alpha_i a_i^\dagger a_i + \beta_\nu b_\nu^\dagger b_\nu + \gamma_{i\nu} b_{e,i\nu}^\dagger b_{e,i\nu} + \epsilon_{i\nu} b_{s,i\nu}^\dagger b_{s,i\nu})}, \quad (\text{A.2})$$

where $\alpha, \beta, \gamma, \epsilon$ are some arbitrary frequencies. The requirement of eliminating all explicit time dependencies in equation (A.1) leads to the conditions

$$\begin{aligned} \gamma_{i\nu} &= \omega_T + \epsilon_{i\nu} \\ \alpha_i &= \gamma_{i\nu} \\ \beta_\nu &= \gamma_{i\nu}. \end{aligned} \quad (\text{A.3})$$

Note that in general, it is not possible to bring the levels $|g\rangle$ and $|s\rangle$ to degeneracy for all atoms since $\omega_s + \omega_T = \omega_i = \omega_\nu$ cannot be satisfied for all \bar{i} and $\bar{\nu}$ at the same time. On the other hand, one has a freedom in the choice of frequencies $\alpha - \epsilon$, provided the conditions (A.3) are satisfied. We adopt the following choice

$$\begin{aligned} \alpha_i &= \beta_\nu = \gamma_{i\nu} = \omega_{\text{aux}} \\ \epsilon_{i\nu} &= \omega_{\text{aux}} - \omega_T, \end{aligned} \quad (\text{A.4})$$

where ω_{aux} is an arbitrary auxiliary frequency. The Hamiltonian (2.1) then becomes

$$\begin{aligned} H = & \Delta_i a_i^\dagger a_i + \Delta_\nu b_\nu^\dagger b_\nu + \Delta_e b_{e,i\nu}^\dagger b_{e,i\nu} + \Delta_s b_{s,i\nu}^\dagger b_{s,i\nu} \\ & + \Omega(b_{e,i\nu}^\dagger b_{s,i\nu} + b_{s,i\nu}^\dagger b_{e,i\nu}) + g(b_{e,i\nu}^\dagger b_{g,i\nu}(a_i + b_\nu) + \text{h.c.}), \end{aligned} \quad (\text{A.5})$$

where $\Delta_x = \omega_x - \omega_{\text{aux}}$, $x = \bar{i}, \bar{\nu}, e$ and $\Delta_s = \omega_s - (\omega_{\text{aux}} - \omega_T)$. This leads to the equations of motion

$$\begin{aligned} i\dot{b}_{e,\bar{i}\bar{\nu}} &= \Delta_e b_{e,\bar{i}\bar{\nu}} + \Omega b_{s,\bar{i}\bar{\nu}} + g b_{g,\bar{i}\bar{\nu}}(a_{\bar{i}} + b_{\bar{\nu}}) \\ i\dot{b}_{g,\bar{i}\bar{\nu}} &= g b_{e,\bar{i}\bar{\nu}}(a_{\bar{i}}^\dagger + b_{\bar{\nu}}^\dagger) \\ i\dot{b}_{s,\bar{i}\bar{\nu}} &= \Delta_s b_{s,\bar{i}\bar{\nu}} + \Omega b_{e,\bar{i}\bar{\nu}} \\ i\dot{a}_{\bar{i}} &= \Delta_i a_{\bar{i}} + g \sum_{\nu} b_{g,\bar{i}\nu}^\dagger b_{e,\bar{i}\nu} \\ i\dot{b}_{\bar{\nu}} &= \Delta_\nu b_{\bar{\nu}} + g \sum_{\bar{i}} b_{g,\bar{i}\bar{\nu}}^\dagger b_{e,\bar{i}\bar{\nu}}, \end{aligned} \quad (\text{A.6})$$

where the summation over ν and i in the last two equations is emphasized. Setting $\dot{b}_{e,i\bar{\nu}} = 0$ and substituting $\dot{b}_{e,i\bar{\nu}}$ to the remaining equations yields

$$\begin{aligned} i\dot{b}_{g,i\bar{\nu}} &= \left(-\frac{g^2}{\Delta_e}\right)(a_i^\dagger + b_{\bar{\nu}}^\dagger)(a_{\bar{i}} + b_{\bar{\nu}})b_{g,i\bar{\nu}} + \left(-\frac{g\Omega}{\Delta_e}\right)(a_i^\dagger + b_{\bar{\nu}}^\dagger)b_{s,i\bar{\nu}} \\ i\dot{b}_{s,i\bar{\nu}} &= \Delta_s b_{s,i\bar{\nu}} + \left(-\frac{\Omega^2}{\Delta_e}\right)b_{s,i\bar{\nu}} + \left(-\frac{g\Omega}{\Delta_e}\right)b_{g,i\bar{\nu}}(a_{\bar{i}} + b_{\bar{\nu}}) \\ i\dot{a}_{\bar{i}} &= \Delta_{\bar{i}} a_{\bar{i}} + \sum_{\nu} \left(-\frac{g^2}{\Delta_e}\right)b_{g,i\bar{\nu}}^\dagger b_{g,i\bar{\nu}}(a_{\bar{i}} + b_{\bar{\nu}}) + \left(-\frac{g\Omega}{\Delta_e}\right)b_{g,i\bar{\nu}}^\dagger b_{s,i\bar{\nu}} \\ i\dot{b}_{\bar{\nu}} &= \Delta_{\bar{\nu}} b_{\bar{\nu}} + \sum_i \left(-\frac{g^2}{\Delta_e}\right)b_{g,i\bar{\nu}}^\dagger b_{g,i\bar{\nu}}(a_i + b_{\bar{\nu}}) + \left(-\frac{g\Omega}{\Delta_e}\right)b_{g,i\bar{\nu}}^\dagger b_{s,i\bar{\nu}}. \end{aligned} \quad (\text{A.7})$$

Switching back to the Pauli matrices representation, now in the $\{|s\rangle, |g\rangle\}$ basis, the effective Hamiltonian reads

$$H = \Delta_i a_i^\dagger a_i + \Delta_{\bar{\nu}} b_{\bar{\nu}}^\dagger b_{\bar{\nu}} + \frac{\tilde{\omega}_{a,i\bar{\nu}}}{2} \sigma_{i\bar{\nu}}^z + \tilde{g} (\sigma_{i\bar{\nu}}^+ (a_i + b_{\bar{\nu}}) + \text{h.c.}) + F_{i\bar{\nu}}, \quad (\text{A.8})$$

where

$$\begin{aligned} \tilde{\omega}_{a,i\bar{\nu}} &= \Delta_s - \frac{\Omega^2 - g^2(a_i^\dagger + b_{\bar{\nu}}^\dagger)(a_{\bar{i}} + b_{\bar{\nu}})}{\Delta_e} \\ \tilde{g} &= -\frac{g\Omega}{\Delta_e} \\ F_{i\bar{\nu}} &= \frac{1}{2} \left(\Delta_s - \frac{\Omega^2 + g^2(a_i^\dagger + b_{\bar{\nu}}^\dagger)(a_{\bar{i}} + b_{\bar{\nu}})}{\Delta_e} \right). \end{aligned} \quad (\text{A.9})$$

The corresponding equations of motion for the operators read

$$i\dot{a}_{\bar{i}} = (\Delta_{\text{ph}}^{\bar{i}} - i\kappa + \lambda\sigma_{i\bar{\nu}}^z) a_{\bar{i}} + \tilde{g}\sigma_{i\bar{\nu}}^- + \lambda b_{\bar{\nu}}(\sigma_{i\bar{\nu}}^z - \mathbb{1}), \quad (\text{A.10a})$$

$$i\dot{b}_{\bar{\nu}} = (\Delta_{\text{ph}}^{\bar{\nu}} - i\kappa + \lambda\sigma_{i\bar{\nu}}^z) b_{\bar{\nu}} + \tilde{g}\sigma_{i\bar{\nu}}^- + \lambda a_i(\sigma_{i\bar{\nu}}^z - \mathbb{1}), \quad (\text{A.10b})$$

$$i\dot{\sigma}_{i\bar{\nu}}^- = \left[\Delta_{\text{at}} - i\frac{\gamma}{2} + 2\lambda(a_i^\dagger + b_{\bar{\nu}}^\dagger)(a_{\bar{i}} + b_{\bar{\nu}}) \right] \sigma_{i\bar{\nu}}^- - \tilde{g}(a_{\bar{i}} + b_{\bar{\nu}})\sigma_{i\bar{\nu}}^z, \quad (\text{A.10c})$$

$$i\dot{\sigma}_{i\bar{\nu}}^z = 2\tilde{g}[\sigma_{i\bar{\nu}}^+(a_{\bar{i}} + b_{\bar{\nu}}) - \text{h.c.}] - i\gamma(\sigma_{i\bar{\nu}}^z + \mathbb{1}), \quad (\text{A.10d})$$

where

$$\begin{aligned} \Delta_{\text{at}} &= \Delta_s - \frac{\Omega^2}{\Delta_e} \\ \Delta_{\text{ph}}^l &= \Delta_l - \frac{g^2}{2\Delta_e} \\ \lambda &= -\frac{g^2}{2\Delta_e} \end{aligned} \quad (\text{A.11})$$

and we have introduced the cavity and spin decays κ , γ which we take to be the same for all cavity modes and all spins respectively.

Appendix B. General stability matrix

The matrix M used in the stability study, equation (3.23), can be simply obtained from the MF equations of motion (3.8) and reads

$$M = \begin{bmatrix} -\kappa & \Delta_{\text{ph}} + \lambda\bar{w} & 0 & -\frac{\tilde{g}}{2} & \lambda\bar{\alpha}_I \\ -\Delta_{\text{ph}} - \lambda\bar{w} & -\kappa & -\frac{\tilde{g}}{2} & 0 & -\lambda\bar{\alpha}_R \\ -4\lambda\bar{\alpha}_R\bar{s}_y & -2\tilde{g}\bar{w} - 4\lambda\bar{\alpha}_I\bar{s}_y & 0 & -(\Delta_{\text{at}} + 2\lambda|\bar{\alpha}|^2) & -2\tilde{g}\bar{\alpha}_I \\ -2\tilde{g}\bar{w} + 4\lambda\bar{\alpha}_R\bar{s}_x & 4\lambda\bar{\alpha}_I\bar{s}_x & (\Delta_{\text{at}} + 2\lambda|\bar{\alpha}|^2) & 0 & -2\tilde{g}\bar{\alpha}_R \\ 2\tilde{g}\bar{s}_y & 2\tilde{g}\bar{s}_x & 2\tilde{g}\bar{\alpha}_I & 2\tilde{g}\bar{\alpha}_R & 0 \end{bmatrix}, \quad (\text{B.1})$$

where $|\bar{\alpha}|^2 = \bar{\alpha}_R^2 + \bar{\alpha}_I^2$ and \bar{v} are the steady state solutions, $\bar{v} \in \{\bar{\alpha}_R, \bar{\alpha}_I, \bar{s}_x, \bar{s}_y, \bar{w}\}$.

Appendix C. Large λ expansion

Here we seek a perturbative solution of the algebraic steady state equations (3.9) and (3.25) in the large λ limit. This can be done using a perturbative ansatz for the variables of the form

$$v = \sum_{n=0}^{\infty} \lambda^{-n} v^{(n)} = v^{(0)} + \lambda^{-1} v^{(1)} + \lambda^{-2} v^{(2)} + \dots \quad (\text{C.1})$$

The set of equations of order λ^1 read

$$\begin{aligned} \alpha_I^{(0)} w^{(0)} &= 0 \\ \alpha_R^{(0)} w^{(0)} &= 0 \\ (\alpha_I^{(0)2} + \alpha_I^{(0)2}) s_y^{(0)} &= 0 \\ (\alpha_I^{(0)2} + \alpha_I^{(0)2}) s_x^{(0)} &= 0, \end{aligned} \quad (\text{C.2})$$

which yield the solution for $\alpha_R^{(0)} = \alpha_I^{(0)} = 0$. Using this result, order λ^0 equations simplify to

$$\alpha_R^{(1)} = -\frac{\tilde{g} s_x^{(0)} + 2\eta_R}{2w^{(0)}}, \quad (\text{C.3a})$$

$$\alpha_I^{(1)} = \frac{\tilde{g} s_y^{(0)} - 2\eta_I}{2w^{(0)}}, \quad (\text{C.3b})$$

$$s_x^{(0)2} + s_y^{(0)2} + w^{(0)2} = N^2. \quad (\text{C.3c})$$

In order to proceed, we realize that the equation (3.25e) at order λ^{-1} reads

$$\alpha_R^{(1)} s_y^{(0)} = -\alpha_I^{(1)} s_x^{(0)}. \quad (\text{C.4})$$

Substituting for $\alpha_R^{(1)}$, $\alpha_I^{(1)}$ from (C.3a) and (C.3b), it can be cast to the form

$$-\frac{s_y^{(0)}}{s_x^{(0)}} = \frac{\eta_I}{\eta_R}. \quad (\text{C.5})$$

In order to simplify the treatment further, we put $\eta_I = 0$ which implies $s_y^{(0)} = 0$ and by the sake of (C.3b) $\alpha_I^{(1)} = 0$. We will use these solutions in what follows.

Order λ^{-1} equations, after the substitutions of the solutions $\alpha_R^{(0)} = \alpha_I^{(0)} = 0$ read

$$0 = -\kappa \alpha_R^{(1)} + (\Delta_{\text{ph}} + w^{(1)}) \alpha_I^{(1)} - \frac{\tilde{g}}{2} s_y^{(1)} + \alpha_I^{(2)} w^{(0)}, \quad (\text{C.6a})$$

$$0 = -(\Delta_{\text{ph}} + w^{(1)}) \alpha_R^{(1)} - \kappa \alpha_I^{(1)} - \frac{\tilde{g}}{2} s_x^{(1)} - \alpha_R^{(2)} w^{(0)}, \quad (\text{C.6b})$$

$$0 = \tilde{g} \alpha_I^{(1)} w^{(0)} + |\alpha^{(1)}|^2 s_y^{(0)}, \quad (\text{C.6c})$$

$$0 = \tilde{g} \alpha_R^{(1)} w^{(0)} - |\alpha^{(1)}|^2 s_x^{(0)}, \quad (\text{C.6d})$$

$$0 = w^{(0)} w^{(1)} + s_x^{(0)} s_x^{(1)} + s_y^{(0)} s_y^{(1)}. \quad (\text{C.6e})$$

When substituting the solutions for $s_y^{(0)} = \alpha_I^{(1)} = 0$, (C.6c) is trivial and the equations (C.6) simplify to

$$0 = -\kappa \alpha_R^{(1)} - \frac{\tilde{g}}{2} s_y^{(1)} + \alpha_I^{(2)} w^{(0)}, \quad (\text{C.7a})$$

$$0 = -(\Delta_{\text{ph}} + w^{(1)}) \alpha_R^{(1)} - \frac{\tilde{g}}{2} s_x^{(1)} - \alpha_R^{(2)} w^{(0)}, \quad (\text{C.7b})$$

$$0 = \tilde{g} \alpha_R^{(1)} w^{(0)} - \alpha_R^{(1)2} s_x^{(0)}, \quad (\text{C.7c})$$

$$0 = w^{(0)} w^{(1)} + s_x^{(0)} s_x^{(1)}. \quad (\text{C.7d})$$

Substituting the solution for $\alpha_R^{(1)}$ from (C.3a) into (C.7c) yields

$$(\tilde{g} s_x^{(0)} + 2\eta_R)[2\tilde{g} w^{(0)2} + (\tilde{g} s_x^{(0)} + 2\eta_R) s_x^{(0)}] = 0 \quad (\text{C.8})$$

which has the solutions

$$s_x^{(0)} = -\frac{2\eta}{\tilde{g}}, \quad (\text{C.9a})$$

$$s_{x,\pm}^{(0)} = \frac{\eta}{\tilde{g}} \left(1 \pm \sqrt{1 + \frac{2\tilde{g}^2 N^2}{\eta^2}} \right), \quad (\text{C.9b})$$

where $\eta = \eta_R$ and we have used the spin conservation $w^{(0)2} + s_x^{(0)2} = N^2$. Since $|s_x^{(0)}| \leq N$, the solution (C.9a) is valid for $\tilde{g} \geq g_1^* = 2\eta/N$. Next, we remark, that the function $s_{x,+}^{(0)}(\tilde{g})$ is monotonously decreasing with the limit

$$\lim_{\tilde{g} \rightarrow \infty} s_{x,+}^{(0)}(\tilde{g}) = \sqrt{2}N > N, \quad (\text{C.10})$$

i.e. is unphysical. We are thus left with the solution $s_{x,-}^{(0)}(\tilde{g})$ which is also monotonously decreasing function with the asymptotes

$$\begin{aligned} \lim_{\tilde{g} \rightarrow 0} s_{x,-}^{(0)}(\tilde{g}) &= 0 \\ \lim_{\tilde{g} \rightarrow \infty} s_{x,-}^{(0)}(\tilde{g}) &= -\sqrt{2}N. \end{aligned} \quad (\text{C.11})$$

It is thus clear that the solution $s_{x,-}^{(0)}$ is valid for $g \in [0, g^*]$, where g^* is some critical value for which $s_{x,-}^{(0)}$ reaches the physically allowed maximum $|s_{x,-}^{(0)}| = N$. It is easy to find, that $g^* = g_1^*$. This completes the leading order solutions and yields the expressions (3.27a) and (3.27b).

The equations (C.2), (C.3) and (C.6) yield a closed set for spin variables up to the order λ^{-1} and for α up to λ^{-2} . It is straightforward to find the solutions up to the respective order explicitly, giving however rather lengthy algebraic expressions. These can be simplified in specific situations. For example, as we discuss in the main text, when looking for solutions in the vicinity of the transition point $g_1^* = 2\eta/N$, the leading order contribution to $|\alpha|^2$ is of order λ^{-2} , namely

$$|\alpha|^2 = \frac{1}{\lambda^2} \alpha_R^{(1)2} + O(\lambda^{-3}), \quad (\text{C.12})$$

which gives the relation (3.28). Similarly, the leading contribution to w is of order λ^0 yielding the expressions (3.29).

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