Natural theory of dark energy

Lorenzo Bordin[®],^{*} Francesc Cunillera[®],[†] Antoine Lehébel[®],[‡] and Antonio Padilla[§] School of Physics and Astronomy, University of Nottingham, Nottingham NG7 2RD, United Kingdom

(Received 20 January 2020; accepted 14 April 2020; published 23 April 2020)

We propose a mechanism that generates a naturally light dark energy field (with Hubble-scale mass), starting from a theory with exclusively high-scale (Planckian) couplings. It is derived from the clockwork model, with $\mathcal{O}(100)$ scalar fields interacting among themselves, as well as with a 4-form field strength. We explicitly embed the key features of our model in type IIA supergravity. We also give an alternative interpretation in a braneworld setup.

DOI: 10.1103/PhysRevD.101.085012

I. INTRODUCTION

The expansion of the Universe is accelerating. There is now strong evidence for this acceleration coming from a slew of cosmological data, mostly notably the observation of distant supernovae [1,2] and measurements of the cosmic microwave background radiation [3]. This acceleration is feeble, driven by dark energy with an energy scale set by the Hubble constant, $H_0 \sim 10^{-33}$ eV [4]. If dark energy corresponds to vacuum energy, its energy density is at least 60 orders of magnitude below its natural value [5-8]. A popular alternative is quintessence, in which dark energy is driven by a dynamical scalar field, mirroring the inflationary dynamics of the early Universe, albeit at a much lower scale. However, if dark energy is truly dynamical, one has to assume that the large vacuum energy is canceled using some other mechanism,¹ so that the current cosmic acceleration is entirely due to the scalar dynamics. Even with this compromise, there is often a further tuning, since the effective mass of the scalar must lie at or below the Hubble scale in order to be cosmologically relevant on the largest scales today. In a typical quintessence model, the light mass is usually put in by hand by assuming a very wide potential, but this is difficult to engineer from a consistent microscopic theory where the typical mass scales are much higher, usually around the Planck scale. Further, as is well known from the electroweak hierarchy problem, light scalar masses are susceptible to large radiative corrections [16].

In this paper, we propose a model of dark energy driven by a pseudoscalar field whose superlight mass emerges naturally from a simple microscopic theory with uniquely high-scale couplings. The model is a dark energy avatar of the pion, its low mass emerging from the spontaneous breaking of a weakly broken symmetry. It consists of $\mathcal{O}(100)$ pseudoscalar fields (axions) with nontrivial mass mixing, one of which has a bilinear mixing with a 4-form field strength. All mass scales in the theory are assumed to lie at, or close to, the Planck scale. We will explicitly demonstrate how the key ingredients can arise naturally in string compactifications.

The model is a marriage of two axion models that have been developed in recent years for different reasons. The first is the clockwork axion [17,18] (see also Refs. [19-32] for related work and interesting applications), proposed in order to account for the super-Planckian decay constants required by models of cosmological relaxation [33]. The basic idea is to have a modest number of axions, ϕ_i , whose masses mix with some characteristic strength q > 1, so that they behave like the gears of a clockwork. If an external source is coupled with some strength to one end of the clockwork, then the resulting low-energy effective field theory (EFT) contains a zero mode that couples to the source with an exponentially weaker coupling. The structures of the mass mixings are crucial. They allow for a nonlinearly realized shift symmetry on the axions $\phi_i \rightarrow \phi_i + cq^{-i}$, where the size of the shift is warped by factors of q as we move through the gears. This symmetry can be identified with an asymmetrically distributed unbroken subgroup of an underlying $U(1)^N$, with the axions identified as pions below the scale of spontaneous symmetry breaking. Further, the shift symmetry guarantees the existence of a zero mode in the low-energy EFT, while the warping ensures that its overlap with the far end of the clockwork is small. This is what suppresses the coupling to external sources.

lorenzo.bordin@nottingham.ac.uk

francesc.cunilleragarcia@nottingham.ac.uk

antoine.lehebel@nottingham.ac.uk

antonio.padilla@nottingham.ac.uk

¹An interesting proposal for eliminating vacuum energy is given by *vacuum energy sequestering* [9–15]. This setup is particularly appealing because it eliminates vacuum energy while allowing for acceleration through a slowly rolling inflaton or quintessence field [10].

Published by the American Physical Society under the terms of the Creative Commons Attribution 4.0 International license. Further distribution of this work must maintain attribution to the author(s) and the published article's title, journal citation, and DOI. Funded by SCOAP³.

The second model in the marriage is the field theory model of axions coupled to 4-forms, originally proposed by Dvali [34,35], then cleverly applied to cosmology by Kaloper and Sorbo [36,37] (see also Ref. [38]). This has been proposed as a field theory realization of axion monodromy in string theory [39,40]. The idea is to introduce a bilinear mixing between the axion and the 4-form field strength. The theory admits a dual description in terms of a massive pseudoscalar and the magnetic flux of the 4-form. The latter is locally constant in spacetime, although it can jump between quantized values across a three-dimensional membrane. The model is of particular interest to single-field inflation, since it avoids some of the problems associated with super-Planckian field excursions for the inflaton. This is because the effective inflaton is a gauge invariant combination of the axion and the magnetic flux-its large field values are obtained through the flux, which in turn may be identified with macroscopic quantities characterizing the system rather than high-scale excitations of the inflaton field [41]. Small deformations of these models can also give rise to an emergent theory of vacuum energy sequestering, screening the effects of vacuum energy at large scales [42–44].

In the Dvali-Kaloper-Sorbo model, the inflaton gets a mass from its mixing with the 4-form field strength, and the stronger the coupling, the larger the effective mass. It is this mechanism for generating a mass that we exploit in our model of dark energy. By coupling one end of our clockwork to the 4-form, we guarantee that the mixing with the zero mode of the clockwork is exponentially small, and as a result, a very small mass is generated in the low-energy effective theory. It should be relatively straightforward to generate the key ingredients of these models from fundamental theory, and we illustrate this with a toy model involving a compactification of type IIA string theory on a Kähler manifold of the form $T^2 \times \Sigma_q^4$, where Σ_q^4 is a manifold of dimension 4, and sufficiently high genus. To further enhance the prospect of deriving this theory as the low-energy limit of some UVcomplete theory, we use the clockwork to deconstruct an extra dimension. This motivates a family of braneworld configurations where the low scale of dark energy emerges naturally on the brane.

The rest of this paper is organized as follows: In the next section, we build the discrete version of our model as a marriage of the two axion proposals described above. In Sec. III, we demonstrate how such a setup can be obtained from a toy compactification of type IIA string theory. In Sec. IV, we perform the dimensional deconstruction of the discrete model, and study a family of braneworld configurations showing how the braneworld obsverver measures a low scale of dark energy on account of the warping of the extra dimension. In Sec. V, we conclude.

II. CLOCKWORK DARK ENERGY

We begin with a chain of N + 1 pseudoscalar fields $\phi_0, ..., \phi_N$, all characterized by a single ultraviolet mass

scale M and a nearest-neighbor interaction with strength q [17,18]. The mass M is assumed to lie at or close to the Planck scale. We further assume that one end of the chain is coupled to a 4-form field strength as in the Dvali-Kaloper-Sorbo model. The combined setup is described by the following Lagrangian density:

$$\mathcal{L} = -\frac{1}{2} \left[\sum_{i=0}^{N} (\partial \phi_i)^2 + M^2 \sum_{i=0}^{N-1} (\phi_i - q\phi_{i+1})^2 \right] + \frac{\mu}{24} \phi_N \frac{\epsilon^{\mu\nu\alpha\beta}}{\sqrt{-g}} F_{\mu\nu\alpha\beta} - \frac{1}{48} F_{\mu\nu\alpha\beta} F^{\mu\nu\alpha\beta}.$$
(1)

In the first line, we recognize the clockwork model of Refs. [17,18]. In principle, we could allow for sitedependent masses $M_i \sim M$ and mixing strengths $q_i \sim q$, although for simplicity we take them all to be equal. The coupling q is dimensionless and assumed to be greater than 1, but it remains of order unity. The second line of Eq. (1)contains the Dvali-Kaloper-Sorbo model for the Nth site in the chain. Here $F_{\mu\nu\rho\sigma} = 4\partial_{[\mu}A_{\nu\alpha\beta]}$ is the 4-form field strength, $\epsilon^{\mu \nu \rho \sigma}$ is the totally antisymmetric Levi-Civita symbol defined such that $e^{0123} = 1$, and indices are raised and lowered with respect to the metric $g_{\mu\nu}$. The bilinear mixing between the axion and the 4-form reveals another mass scale, $\mu \sim M$, which we also assume to be given by the characteristic ultraviolet scale of the theory. The gravitational sector of the theory is assumed to be described by Einstein gravity, although we will not need to include that in our discussion.

It is well known that the clockwork Lagrangian gives rise to a massless Goldstone pseudoscalar [17-19], a consequence of the nonlinearly realized shift symmetry $\phi_i \rightarrow$ $\phi_i + cq^{-i}$ for arbitrary c. This symmetry remains perturbatively unbroken through the mixing with the 4-form, although in the presence of charged membranes, nonperturbative effects break the continuous symmetry down to a discrete subgroup [45,46]. However, the symmetry is only *mildly* broken, because ϕ_N has exponentially suppressed overlap with the zero mode. We thus expect that the zero mode acquires a mass, but that the latter remains tiny. In order to see this explicitly, it is convenient to integrate out the 4-form and pass to a dual description in which the 4-form mixing generates a new mass term for the last axion in the chain. This can be done in a straightforward manner by adding a Lagrange multiplier term of the form $\frac{1}{24}Q \frac{\epsilon^{\mu\nu\alpha\beta}}{\sqrt{-q}} (F_{\mu\nu\rho\sigma} - 4\partial_{[\mu}A_{\nu\rho\sigma]})$, then eliminating F using its algebraic equation of motion² [36]. This yields a dual theory described by the following Lagrangian:

²Note that since the theory is quadratic in F, this amounts to performing the Gaussian in the path integral, which can, of course, be done exactly.

$$\mathcal{L} = -\frac{1}{2} \left[\sum_{i=0}^{N} (\partial \phi_i)^2 + M^2 \sum_{i=0}^{N-1} (\phi_i - q\phi_{i+1})^2 \right] \\ -\frac{1}{2} (Q + \mu \phi_N)^2 - \frac{Q}{6} \frac{\epsilon^{\mu\nu\alpha\beta}}{\sqrt{-g}} \partial_{[\mu} A_{\nu\alpha\beta]}.$$
(2)

The Lagrange multiplier Q is fixed to be constant by the variation of the 3-form. If the latter is coupled to membrane

$$\mathcal{M}_{ij} = M^2 \begin{bmatrix} 1 & -q & 0 \\ -q & 1+q^2 & -q \\ 0 & -q & 1+q^2 \\ & \ddots \\ \vdots & & \ddots \\ 0 & & \cdots \end{bmatrix}$$

charges, then Q is quantized in units of the membrane charge, e, as in $\langle Q \rangle = 2\pi \mathcal{N}_Q e$ for integer values of \mathcal{N}_Q [45,46]. This quantization condition is compatible with the unbroken symmetry transformations, which take the form $\phi_i \to \phi_i + 2\pi n \frac{e}{\mu} q^{N-i}$, $Q \to Q + 2\pi n e$ for integer values of n. The mass matrix in the dual description is given by

where $r = (\mu/M)^2$ is the square of the ratio between the two mass scales. The eigenmasses of this matrix are given by the roots of an (N + 1)th-order polynomial. It turns out that there is a tower of N massive modes whose masses go with the ultraviolet scale M. The remaining mode is massless in the limit where $r \rightarrow 0$, and in general is ultralight. We can find it by linearizing the above eigenvalue problem; the resulting ultralight mass scale is given by

$$m_0^2 \simeq \frac{(q^2 - 1)^2 r M^2}{q^{2(N+1)}(q^2 + r - 1) + (N+1)r(1 - q^2) - q^2 - r + 1}.$$
(4)

After integrating out the 4-form and setting the corresponding flux to zero, the emergent potential for the ultralight mode, π_0 , is given by a simple quadratic:

$$V_0 = \frac{1}{2} m_0^2 \pi_0^2. \tag{5}$$

In the limit of large N (and for q larger than unity), m_0 is well approximated by

$$m_0^2 \simeq \frac{1}{q^{2(N+1)}} \frac{(q^2 - 1)^2 r M^2}{(q^2 + r - 1)}.$$
 (6)

Thus, m_0 acquires an exponential suppression with respect to M, the argument in the exponential being the total number of clockwork gears, N + 1. If we take $M = \mu = m_{\rm P}$ $(m_{\rm P} = 1.22 \times 10^{19} \text{ GeV} \text{ being the Planck mass}), q = 2,$ and N = 200, we get $m_0 \simeq 5.7 \times 10^{-33}$ eV, which is the energy scale associated with the Hubble rate.

At large scales, the dynamics will be equivalent to quintessence driven by a quadratic potential. However, the mass scale of potential has not been tuned to the tiny value demanded by nature; rather, it has arisen naturally on account of the clockwork mechanism and the coupling to the 4-form. The underlying theory is made up uniquely of high-scale couplings. We emphasize the fact that the mass mixings need not be identical but can have site dependence. As long as they are greater than unity, the clockwork mechanism will kick in as usual, even for order-1 couplings, and the suppression of the mass of the ultralight mode will occur as desired.

Finally, we note that if we allowed for a nonvanishing flux, the potential would actually go as

$$V_0 = \frac{1}{2} m_0^2 \left(\frac{2\pi \mathcal{N}_Q e}{m_0} + \pi_0 \right)^2.$$
(7)

For natural values of $\pi_0 \sim m_P$, the flux term dominates the potential for nonvanishing \mathcal{N}_Q , and there is too much dark energy. This is why we assumed a vanishing flux $\mathcal{N}_{O} = 0$, which is a robust condition provided the nucleation rate for bubbles of nonvanishing flux is suppressed.³

³Although a detailed analysis of this interesting question is beyond the scope of our work, we expect that there is indeed suppression for the following reason: For $e \sim m_P^2$, a unit of flux generates a large positive Planckian potential. Therefore, we can crudely model the transition from vanishing to nonvanishing flux in terms of a bubble of Planckian de Sitter curvature nucleating in a quasi-de Sitter background with a very small curvature. Such processes, tunneling from true to false, are possible but are known to be suppressed relative to those going from false to true, becoming infinitely suppressed in the limit that the low-scale curvature approaches zero (see, e.g., Ref. [47]).

III. CLOCKWORK DARK ENERGY FROM STRING THEORY

Our model can be motivated from the point of view of a higher-dimensional theory. Indeed, in a theory with extra dimensions, a large number of scalar fields in the 4D EFT is often associated with the periods of differential p-forms living in the higher-dimensional theory. In Refs. [48,49], the authors showed that one may rewrite the democratic type IIA supergravity formulation [50] in terms of a pseudoaction containing Minkowski 4-forms and its dual fields, which is equivalent to the democratic action at the level of the equations of motion.

The democratic action is usually written in terms of a polyform gauge invariant field strength, $G = G_0 + G_2 + \cdots + G_{10}$, where $\mathbf{G} = d\mathbf{C} + B \wedge \mathbf{C} + \mathcal{F} \wedge e^B$ and *B* is the Kalb-Ramond 2-form, $\mathbf{C} = C_1 + C_3 + \cdots + C_9$ is the polyform gauge field, and $\mathcal{F} = \mathcal{F}_0 + \mathcal{F}_2 + \cdots + \mathcal{F}_{10}$ is a formal sum of internal fluxes only. Let us take the following ansatz for the nonvanishing supergravity fields:

$$B_2 = b^i(x)\omega_i, \qquad C_3 = c_3^0(x), \qquad C_5 = c_3^i(x)\omega_i,$$

$$C_7 = \bar{c}_{3i}(x)\bar{\omega}^i, \qquad \mathcal{F}_2 = q^i\omega_i, \qquad \mathcal{F}_4 = e_i\bar{\omega}^i, \qquad (8)$$

where we note that this corresponds to the massless limit of type IIA supergravity. We have introduced the cohomology basis of $\{2, 4\}$ -forms in the internal manifold \mathcal{M}_6 as $\{\omega_i, \bar{\omega}^i\}$, respectively, and ω_6 will denote the volume form of \mathcal{M}_6 . With this ansatz, **G** reads⁴

$$G_{2} = q^{i}\omega_{i}, \qquad G_{4} = F_{4}^{0} + (e_{i} + \mathcal{K}_{ijk}b^{j}q^{k})\bar{\omega}^{i},$$

$$G_{6} = F_{4}^{i}\omega_{i} + \left(e_{i}b^{i} + \frac{1}{2}\mathcal{K}_{ijk}b^{j}b^{j}q^{k}\right)\omega_{6}, \qquad G_{8} = \bar{F}_{4i}\bar{\omega}^{i},$$

$$(9)$$

with $\mathcal{K}_{ijk} = \int_{\mathcal{M}_6} \omega_i \wedge \omega_j \wedge \omega_k$, the triple intersection numbers of \mathcal{M}_6 .

Below we describe the effective 4D action; in doing so, we will assume that the volume moduli, including the dilaton, are stabilized by some other ingredient of the theory (see, e.g., Ref. [51]). Our main objective is to obtain the main characteristic of the clockwork mechanism after compactifying the theory down to four dimensions.

These fields enter the 4D potential as follows (see for details Refs. [48,49]):

$$V_{4D} \propto F_4^0 \wedge *F_4^0 + g_{ij}F_4^i \wedge *F_4^j + g^{ij}\bar{F}_{4i} \wedge *\bar{F}_{4j} + F_4^0\rho_0 + F_4^i\rho_i + \bar{F}_{4i}\bar{\rho}^i,$$
(10)

⁴We have chosen to use the compact notation

$$F_4^0 = dc_3^0, \qquad F_4^i = dc_3^i + b^i F_4^0, \qquad \bar{F}_{4i} = d\bar{c}_{3i} + \mathcal{K}_{ijk} b^j dc_3^k.$$

where $g_{ij} = \int_{\mathcal{M}_6} \omega_i \wedge \star \omega_j$, $g^{ij} = \int_{\mathcal{M}_6} \bar{\omega}^i \wedge \star \bar{\omega}^j$, and we have introduced the dual scalars to the Minkowski 4-forms:

$$l_s
ho_0 = e_i b^i + rac{1}{2} \mathcal{K}_{ijk} q^i b^j b^k,$$

 $-l_s
ho_i = e_i + \mathcal{K}_{ijk} q^j b^k, \qquad l_s \bar{
ho}^i = q^i.$

Eliminating F_4^i and \bar{F}_{4i} through their equations of motion in favor of their dual scalars, the 4D Lagrangian density with only one Minkowski 4-form is given by

$$\mathcal{L}_{4D} = -\frac{1}{\kappa_4^2} \left[\frac{e^{-2\phi}}{4} g_{ij} db^i \wedge *db^j + \frac{\hat{V}_6^3 e^{\frac{\phi}{2}}}{32} F_4^0 \wedge *F_4^0 -\frac{1}{4} F_4^0 \rho_0 + \frac{l_s^6 e^{\frac{\phi}{2}}}{2\hat{V}_6^2} (g^{ij} \rho_i \rho_j + e^{\phi} g_{ij} \bar{\rho}^i \bar{\rho}^j) \right],$$
(11)

with $\operatorname{Vol}(\mathcal{M}_6) = l_s^6 \hat{V}_6$.

As this stands, the vacuum expectation value for the axions is not necessarily vanishing, owing to the presence of a tadpole in the potential for b_i . We can fix this by shifting $b_i \rightarrow \beta_i + b_i$ for some constant flux β_i chosen so that it eliminates the tadpole. By comparison to Eq. (1), one finds a Lagrangian density of the form⁵

$$\mathcal{L} = -\frac{1}{2} [\gamma_{ij} \partial b^i \partial b^j + M_{ij} b^i b^j + \frac{\alpha}{48} (F_4^0)^2] + \frac{1}{24} \chi_i b^i F_4^0,$$
(12)

with

$$\begin{split} \gamma_{ij} &= \frac{e^{-2\phi}}{2\kappa_4^2} \int_{\mathcal{M}_6} \omega_i \wedge \star \omega_j, \qquad \alpha = \frac{2\hat{V}_6^3 e^{\frac{\phi}{2}}}{3\kappa_4^2}, \\ M_{ij} &= \frac{e^{\frac{\phi}{2}} l_s^4}{\hat{V}_6^2 \kappa_4^2} \mathcal{K}_{ikl} \mathcal{K}_{jmn} q^k q^m \int_{\mathcal{M}_6} \bar{\omega}^l \wedge \star \bar{\omega}^n, \\ \chi_i &= \frac{6}{\kappa_4^2 l_s} (e_i + \mathcal{K}_{ijk} \beta^j q^k). \end{split}$$

The main ingredients of the clockwork dark energy model that we have considered so far are a mass matrix with vanishing eigenvalue and a zero mode that overlaps with the remaining 4-form. The former is the clockwork condition, while the latter allows the zero mode to acquire a small mass. Note that our goal here is to motivate these two key ingredients, rather than the precise form of the clockwork Lagrangian presented in the last section. In any event, if s^i is an eigenvector of M_{ij} with null eigenvalue, then the

⁵In writing Eq. (12), we have dropped a next-to-leading-order correction term to the coupling $(\sim F_4^0 b^2)$ as well as a cosmological constant term and a total derivative that do not contribute to our discussion.

first condition requires $M_{ij}s^i \sim \mathcal{K}_{ikl}q^ks^i = 0$, while the second gives $s^i\chi_i \neq 0$. To comply with these two requisites, we must therefore impose the following condition on the internal geometry:

$$\int_{\mathcal{M}_6} (\omega_i s^i) \wedge \omega_j \wedge (\omega_k q^k) = 0.$$
 (13)

A simple solution to the above constraint is provided by $s^i = q^i$ and the corresponding 2-form given by a product of 1-forms, i.e., $\omega_k q^k = u \wedge v$. In that case, the constraint is immediately satisfied due to the antisymmetry of the wedge product on odd forms. As long as we require $q^i e_i \neq 0$, the second condition can also be satisfied.

The proposed solution could be realized on a Calabi-Yau three-fold of sufficiently high genus. However, a simplified way to visualize the right configuration is given in the presence of a Kähler internal manifold of the form $\mathcal{M}_6 \simeq T^2 \times \Sigma_q^4$, where g indicates the genus of the twodimensional complex surface. Indeed, with this setup one naturally selects the zero mode, q^i , to point along the toroidal directions T^2 . This is because the top form along T^2 is automatically a 2-form, corresponding to a product of the 1-forms of the torus. As long as we take the e_i to overlap slightly with the T^2 , but not completely, we also satisfy the second condition, allowing the zero mode to gain a small mass through its 4-form mixing. Note that an explicit solution to the Einstein equations for a generic Kähler internal manifold that is sufficiently small and compact to be compatible with standard four-dimensional observations may be difficult to find in practice, and we present it here just for illustration. We leave a more detailed investigation to future study.

Finally, we have shown that for the clockwork mechanism to bring the mass of the lightest mode down to the dark energy scale, one requires $\mathcal{O}(200)$ scalar fields. For simplicity of exposition, let us think of these fields as descending only from the Kalb-Ramond 2-form. Then, the previous requirement translates into a constraint on the cohomology structure of Σ_g^4 after fixing the T^2 geometry, namely $g = b_1 - b_2/2$, where b_p is the *p*th Betti number of Σ_q^4 . Using $b_2 \sim \mathcal{O}(200)$, we find that $g \sim b_1 - 100$.

IV. DECONSTRUCTING CLOCKWORK DARK ENERGY

If the number of pseudoscalar fields is very high, $N \rightarrow \infty$, the clockwork mechanism can arise as a deconstruction of an extra compact dimension [19]. In this limit, the clockwork gears merge into a single field Φ , and the gear ϕ_i becomes the value of Φ at site *i*; the interaction of the 4-form with the last site of the discrete clockwork corresponds to the localization of the 4-form on a brane at the boundary of the compactified extra dimension. Our

model can be obtained from the following 5D theory, defined on a fixed geometrical background:

$$S_{5D} = S_{\text{bulk}} + S_0 + S_{\pi R},$$
 (14)

where we have a canonical scalar in the bulk

$$S_{\text{bulk}} = \int d^4x \int_0^{\pi R} dy \sqrt{-g} \left[-\frac{m}{2} g^{IJ} (\partial_I \Phi) (\partial_J \Phi) \right], \quad (15)$$

and two branes, one containing the dark energy sector

$$S_{0} = \int_{y=0} d^{4}x \sqrt{-\gamma_{0}} \left(\frac{\mu}{24} \Phi \frac{\epsilon^{\mu\nu\alpha\beta}}{\sqrt{-\gamma_{0}}} F_{\mu\nu\alpha\beta} - \frac{1}{48} F^{\mu\nu\alpha\beta} F_{\mu\nu\alpha\beta} \right) \bigg],$$
(16)

and the other containing the matter sector

$$S_{\pi R} = \int_{y=\pi R} d^4 x \sqrt{-\gamma_R} \mathcal{L}_m(\gamma_R^{\mu\nu}, \Psi).$$
(17)

We define g_{IJ} to be the 5D metric, with indices I, J running over the four spacetime dimensions x^{μ} , as well as the additional fifth dimension y. For its part, the y dimension extends from 0 to πR , with a reflection symmetry at y = 0and $y = \pi R$. These two boundary surfaces represent the location of the branes with the induced metric on the brane at y = 0, πR given by $\gamma^{0}_{\mu\nu}$, $\gamma^{R}_{\mu\nu}$, respectively. Upon dividing the y dimension into N + 1 sites and discretizing the action (14) accordingly, one can recover the Lagrangian density [Eq. (1)], provided we input the following geometry:

$$ds^{2} = e^{\frac{4ky}{3}}(dy^{2} + dx^{2}), \qquad (18)$$

where dx^2 entails the 4D metric on a brane, and identifying $q^N = e^{\pi kR}$ with the mass scale $M = N/(\pi R)$. We imagine all such scales—M, m, μ , and k—to be of the same order, corresponding to the UV scale of the underlying braneworld theory. For this particular geometry, the latter corresponds to a five-dimensional linear dilaton model with boundary terms living on the two branes [21].

However, the mechanism that suppresses the mass of the lightest mode is very general and works also for metrics that differ from Eq. (18). For instance, it works also in the following family of metrics:

$$ds^{2} = e^{\frac{4ky}{3}}(e^{-4\ell ky}dy^{2} + dx^{2}).$$
 (19)

The clockwork geometry is recovered by choosing $\ell = 0$, while for $\ell = 1/3$ one gets Randall-Sundrum [52]. The equation of motion for Φ , once *F* has been integrated out, is

$$e^{2(\ell-\frac{1}{3})ky}\left[\Phi''+2(1+\ell)k\Phi'+\frac{\Box\Phi}{e^{4k\ell y}}\right] = \delta(y)\frac{\mu(\mu\Phi+Q)}{m},$$
(20)

where a prime stands for a y derivative, and \Box is the 4D d'Alembertian. Setting $\Phi = -Q/\mu + \delta \Phi$, we can easily solve for $\delta \Phi$ in the bulk by taking the 4D Fourier transform of Eq. (20).⁶ The solution is then given by

$$\delta\Phi(y,x^{\mu}) = \int d^4p \left[A(p^2) J_{\frac{1+\ell}{2\ell}} \left(\frac{\sqrt{-p^2}}{2k\ell e^{2k\ell y}} \right) + B(p^2) J_{-\frac{1+\ell}{2\ell}} \left(\frac{\sqrt{-p^2}}{2k\ell e^{2k\ell y}} \right) \right] e^{ip_{\mu}x^{\mu} - k(1+\ell)y},\tag{21}$$

where the $J_{\alpha}(z)$ are Bessel functions of the first kind, *A* and *B* being free functions. Equation (20) further imposes some boundary conditions at y = 0 and πR . As a result, *A* and *B* are linearly related, while p^2 can only take some values among a quantized set, $p^2 = -m_n^2$, $n \in \mathbb{N}$. The masses m_n are found as solutions of the following equation:

$$J_{\frac{\ell-1}{2\ell}}\left(\frac{m_n}{2k\ell}e^{-2k\ell\pi R}\right) \left[J_{\frac{\ell+1}{2\ell}}\left(\frac{m_n}{2k\ell}\right) \left[4k(\ell+1)m+\mu^2\right] - 2m_n m J_{\frac{3\ell+1}{2\ell}}\left(\frac{m_n}{2k\ell}\right)\right] + J_{\frac{1-\ell}{2\ell}}\left(\frac{m_n}{2k\ell}e^{-2k\ell\pi R}\right) \left[\mu^2 J_{-\frac{\ell+1}{2\ell}}\left(\frac{m_n}{2k\ell}\right) - 2m_n m J_{\frac{\ell-1}{2\ell}}\left(\frac{m_n}{2k\ell}\right)\right] = 0.$$

$$(22)$$

They are all of order k, except for the first, which is very light, being suppressed by $e^{k\pi R} \gg 1$. It is possible to find an approximate expression for this light mass:

$$m_0^2(\ell) \simeq \frac{\mu^2}{m} \frac{8k^2\ell(1-\ell^2)}{2e^{2k(1-\ell)\pi R}\ell[2k(1+\ell) + \frac{\mu^2}{m}] - (1+\ell)(4k\ell + \frac{\mu^2}{m}) - e^{-4k\ell\pi R}(\ell-1)\frac{\mu^2}{m}}.$$
(23)

The above approximation is based on an expansion of the Bessel functions, assuming $m_0 \ll k\ell$. This is no longer true when $\ell \rightarrow 0$. To compute the lightest mass in the clockwork geometry, one can solve directly

$$m_0^2(0) = \frac{4k^2\mu^2}{e^{2\pi kR}(4km+\mu^2) - 4km - 2\pi k\mu^2 R + \mu^2}$$

Equation (23) shows that the suppression mechanism works for the whole family of metrics (19), even if it gets less efficient as we depart from the clockwork geometry. As an example, for the values given below Eq. (6), and setting $\mu^2/m = M$, one gets $m_0(\ell = 1/3) \simeq 6.6 \times 10^{-16}$ eV (Randall-Sundrum geometry), while $m_0(\ell = 0) \simeq 6.3 \times$ 10^{-33} eV. Note that Eq. (23) should not be trusted physically when $l \gtrsim 1$, because the length of the *y* dimension drops below the UV scale of the five-dimensional theory.

As it stands, the mass of the lightest dark energy mode is suppressed relative to UV scales set by matter resident on *either* of the two branes. Why, then, have we placed matter on the brane at $y = \pi R$? This is because the energy density of the dark energy field in slow roll is enhanced by the effective four-dimensional Planck scale. This enhancement exactly compensates for the suppression in the mass of the field, and in the end the energy density during slow roll scales like k^4 . We shall postpone further details to a future publication [53]. In any event, it turns out that the energy density of dark energy will only be suppressed if we calibrate our scales relative to the πR brane, where the warp factor is exponentially large. This is why we put the visible sector of our theory on the right-hand brane.

V. CONCLUSIONS

In this paper, we have considered the interplay between clockwork gravity and the coupling between its pseudoscalar fields and Minkowski top-forms. The result of this added coupling is a natural realization of superlight dark energy whose energy scale arises solely from the characteristics of the high-energy theory.

Furthermore, we have considered two possible UV toy models for the EFT. First, we consider a type IIA compactification. Upon constraining the geometry of the internal space, we find a solution that ensures that our toy models will contain a zero eigenvalue on its mass matrix, the key feature of the clockwork mechanism. Second, we propose a generalization of Ref. [19] which contains the desired coupling to a single 4-form living on a brane.

It would be interesting to explore the phenomenology of our natural dark energy scenario in greater detail in each of its different guises. In two particularly exciting developments, work is currently underway to develop the toy supergravity setup to tackle the coincidence problem, or to ease the tension between measurements of the current Hubble scale.

⁶In general, one should embed the action (14) in some geometric theory and study the perturbations of the geometric quantities as well as the fluctuations of the clockwork field. However, these fluctuations decouple at the linear level. We can thus consistently examine the fluctuations of Φ on their own.

ACKNOWLEDGMENTS

A. P. is funded by a Leverhulme Trust Research Project Grant. L. B., A. L., and A. P. acknowledge support from STFC Grant No. ST/P000703/1. F. C. is funded by a studentship at the University of Nottingham.

- S. Perlmutter *et al.* (Supernova Cosmology Project Collaboration), Astrophys. J. 517, 565 (1999).
- [2] A. G. Riess *et al.* (Supernova Search Team), Astron. J. 116, 1009 (1998).
- [3] N. Aghanim *et al.* (Planck Collaboration), arXiv:1807 .06209.
- [4] E. J. Copeland, M. Sami, and S. Tsujikawa, Int. J. Mod. Phys. D 15, 1753 (2006).
- [5] S. Weinberg, Rev. Mod. Phys. 61, 1 (1989).
- [6] J. Polchinski, arXiv:hep-th/0603249.
- [7] C. P. Burgess, arXiv:1309.4133.
- [8] A. Padilla, arXiv:1502.05296.
- [9] N. Kaloper and A. Padilla, Phys. Rev. Lett. 112, 091304 (2014).
- [10] N. Kaloper and A. Padilla, Phys. Rev. D 90, 084023 (2014);
 90, 109901(A) (2014).
- [11] N. Kaloper and A. Padilla, Phys. Rev. Lett. 114, 101302 (2015).
- [12] N. Kaloper, A. Padilla, D. Stefanyszyn, and G. Zahariade, Phys. Rev. Lett. **116**, 051302 (2016).
- [13] N. Kaloper, A. Padilla, and D. Stefanyszyn, Phys. Rev. D 94, 025022 (2016).
- [14] N. Kaloper and A. Padilla, Phys. Rev. Lett. 118, 061303 (2017).
- [15] G. D'Amico, N. Kaloper, A. Padilla, D. Stefanyszyn, A. Westphal, and G. Zahariade, J. High Energy Phys. 09 (2017) 074.
- [16] G. F. Giudice, in *Perspectives on LHC Physics*, edited by G. Kane and A. Pierce (World Scientific, Singapore, 2008), pp. 155–178.
- [17] K. Choi and S. H. Im, J. High Energy Phys. 01 (2016) 149.
- [18] D. E. Kaplan and R. Rattazzi, Phys. Rev. D 93, 085007 (2016).
- [19] G. F. Giudice and M. McCullough, J. High Energy Phys. 02 (2017) 036.
- [20] N. Craig, I. Garcia Garcia, and D. Sutherland, J. High Energy Phys. 10 (2017) 018.
- [21] G.F. Giudice and M. McCullough, arXiv:1705.10162.
- [22] G. F. Giudice, Y. Kats, M. McCullough, R. Torre, and A. Urbano, J. High Energy Phys. 06 (2018) 009.
- [23] D. Teresi, Phys. Lett. B 783, 1 (2018).
- [24] A. Ahmed and B. M. Dillon, Phys. Rev. D 96, 115031 (2017).
- [25] T. Hambye, D. Teresi, and M. H. G. Tytgat, J. High Energy Phys. 07 (2017) 047.
- [26] F. Niedermann, A. Padilla, and P. M. Saffin, Phys. Rev. D 98, 104014 (2018).

- [27] R. Torre, arXiv:1806.04483.
- [28] P. Agrawal, J. Fan, and M. Reece, J. High Energy Phys. 10 (2018) 193.
- [29] A. Goudelis, K. A. Mohan, and D. Sengupta, J. High Energy Phys. 10 (2018) 014.
- [30] S. H. Im, H. P. Nilles, and M. Olechowski, J. High Energy Phys. 01 (2019) 151.
- [31] S. H. Im, H. P. Nilles, and M. Olechowski, J. High Energy Phys. 10 (2019) 159.
- [32] K. Enqvist, S. Hannestad, and M. S. Sloth, Phys. Rev. Lett. 99, 031301 (2007).
- [33] P. W. Graham, D. E. Kaplan, and S. Rajendran, Phys. Rev. Lett. 115, 221801 (2015).
- [34] G. Dvali, Phys. Rev. D 74, 025018 (2006).
- [35] G. Dvali, arXiv:hep-th/0507215.
- [36] N. Kaloper and L. Sorbo, Phys. Rev. Lett. 102, 121301 (2009).
- [37] N. Kaloper and L. Sorbo, Phys. Rev. D 79, 043528 (2009).
- [38] N. Kaloper, A. Lawrence, and L. Sorbo, J. Cosmol Astropart. Phys. 03 (2011) 023.
- [39] E. Silverstein and A. Westphal, Phys. Rev. D 78, 106003 (2008).
- [40] L. McAllister, E. Silverstein, and A. Westphal, Phys. Rev. D 82, 046003 (2010).
- [41] N. Kaloper and A. Lawrence, Phys. Rev. D 95, 063526 (2017).
- [42] A. Padilla, J. High Energy Phys. 01 (2019) 175.
- [43] B. K. El-Menoufi, S. Nagy, F. Niedermann, and A. Padilla, Classical Quantum Gravity 36, 215014 (2019).
- [44] N. Kaloper, J. High Energy Phys. 11 (2019) 106.
- [45] J. Polchinski and A. Strominger, Phys. Lett. B 388, 736 (1996).
- [46] R. Bousso and J. Polchinski, J. High Energy Phys. 06 (2000) 006.
- [47] K. Lee and E. J. Weinberg, Phys. Rev. D 36, 1088 (1987).
- [48] S. Bielleman, L. E. Ibanez, and I. Valenzuela, J. High Energy Phys. 12 (2015) 119.
- [49] F. Carta, F. Marchesano, W. Staessens, and G. Zoccarato, J. High Energy Phys. 09 (2016) 062.
- [50] E. Bergshoeff, R. Kallosh, T. Ortin, D. Roest, and A. Van Proeyen, Classical Quantum Gravity 18, 3359 (2001).
- [51] A. Saltman and E. Silverstein, J. High Energy Phys. 01 (2006) 139.
- [52] L. Randall and R. Sundrum, Phys. Rev. Lett. 83, 3370 (1999).
- [53] L. Bordin, F. Cunillera, A. Lehébel, and A. Padilla (to be published).