1	Paleoclimate change in Ethiopia around the last interglacial derived from
2	annually-resolved stalagmite evidence
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4	Asfawossen Asrat ^{a, *} , Andy Baker ^b , Melanie J. Leng ^c , John Hellstrom ^d , Gregoire Mariethoz ^e ,
5	Ian Boomer ^f , Dorothy Yu ^g , Catherine N. Jex ^h , John Gunn ^f
6	
7	^a School of Earth Sciences Addis Ababa University, P. O. Box 1176, Addis Ababa, Ethiopia
8	^b PANGEA Research Centre, School of Biological, Earth and Environmental Sciences,
9	UNSW Sydney, Sydney, NSW 2052, Australia; and ARC Centre of Excellence for Australian
10	Biodiversity and Heritage
11	^c NERC Isotope Geosciences Facility, British Geological Survey, Keyworth, UK; and School
12	of Biosciences, University of Nottingham, UK
13	^d School of Earth Sciences, University of Melbourne, Parkville, VIC 3052, Australia
14	^e Institute of Earth Surface Dynamics, University of Lausanne, CH-1015, Switzerland
15	^f School of Geography, Earth and Environmental Sciences, University of Birmingham,
16	Edgbaston, Birmingham, UK
17	^g Mark Wainwright Analytical Centre, UNSW Sydney, Sydney 2052, Australia
18	^h ScienceNordic, Carl Jacobsens Vej 16, 2500 Valby, Copenhagen, Denmark
19	
20	*Corresponding author: asfawossen.asrat@aau.edu.et (A. Asrat)
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22	ABSTRACT
23	Oxygen and carbon ($\delta^{18}O/\delta^{13}C$) isotope, growth rate and trace element data are reported for a
24	U-Th dated, annually-laminated stalagmite, GM1 from Goda Mea Cave, Ethiopia. The
25	stalagmite grew intermittently around the last interglacial. The proxy records are used to

26 develop a conceptual growth model of the stalagmite and to assess its potential for revealing 27 a climate signal in this climatically sensitive northeastern African region during an important period in the evolution of Homo sapiens and dispersal of Anatomically Modern Humans out 28 29 of Africa. Speleothem deposition is of short-duration occurring at ~129 ka, ~120 ka, in an undated growth phase, and at ~108 ka; probably due to tectonic activity. δ^{18} O composition is 30 very stable within growth phases (1σ variability < 0.76‰), as are Mg/Ca, Sr/Ca and Ba/Ca, 31 all indicative of well-mixed source-waters. A shift to positive δ^{18} O values and increased 32 33 variability in Mg/Ca, Sr/Ca and Ba/Ca prior to growth hiatuses is observed, indicating a loss of the well-mixed water source prior to growth cessation. Mean δ^{18} O composition (-3.82 to -34 35 7.77‰) is lower than published modern and Holocene stalagmites from the region. Geochemical data, statistical analyses, and a conceptual model of stalagmite growth, 36 demonstrate that climatic conditions recorded by GM1 were wetter than the Holocene. The 37 ~129 ka growth phase particularly presents an annual record of the relative Intertropical 38 Convergence Zone (ITCZ) position. The GM1 record, the oldest high-resolution continental 39 40 climate record from Ethiopia so far published, presents evidence that any early human migrations which occurred during MIS 5 are likely to have occurred during a wet event in 41 northeast Africa. 42

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Key words: Last interglacial; Northeast Africa; speleothem; oxygen isotopes; paleoclimate

46 **1. Introduction**

In Ethiopia, stalagmites provide high-resolution records of past climate and environment
(Asrat et al., 2007; Baker et al., 2007; 2010). Fast-growing, annually-laminated stalagmites
are ubiquitous, due to the strong seasonality of rainfall and the water balance in Ethiopia.
Regular laminae, visible in hand section, can provide precise annual chronology. Annual

51 growth rates of these stalagmites, determined from the thickness of an annual lamina, is at the 52 upper range of those observed in stalagmites (typically ~ 0.5 mm/yr). This is due to the optimal climatic conditions (high temperature and rainfall) for limestone dissolution and re-53 54 precipitation. This rapid growth facilitates the high-resolution sampling of stalagmite calcite. In Ethiopia, the real advantage of using speleothems to provide a paleoclimate proxy 55 record is that they contain information on past rainfall variability in the region. Several major 56 air streams and convergence zones influence the current climate pattern in northeast Africa, 57 whose effects are often compounded by such regional factors as topography and the 58 59 proximity to the oceans (e.g., Nicholson, 1996). The relatively dry north-easterly and southeasterly monsoons and the humid and moisture-laden (rainfall generating), westerly and 60 61 south-westerly air flow of the Congo air stream, generally dominate the regional wind and 62 pressure patterns. The Intertropical Convergence Zone (ITCZ) and the Congo Air Boundary (CAB) separate these major air streams. The passage of the ITCZ (Fig. 1a) dominantly 63 determines the rainy seasons in Ethiopia, while the topography (highland barriers separated 64 65 by a rift zone) modulates the local rainfall distribution. Accordingly, Ethiopian climate has two rainy seasons, one from the northward passage of the ITCZ, called locally the 'big rains' 66 67 (between June and September), which is reliable and whose maxima migrates with the position of the ITCZ. A second rainy season, the 'small rains', is less consistent and occurs 68 69 between March and May with maxima in April. Dryland farming, including subsistence 70 farming, leads to a high dependency on rains in both seasons. Failure of the 'small' rains is common and has occurred in recent years in 2013/2014 and 2015/2016, particularly in the 71 southeastern Ethiopian lowlands bordering the current study area. The climate dynamical 72 73 cause of the failure of the 'small' rains, and how this varies over time, is still poorly understood. 74

75 In addition, reliable, high-resolution climate records beyond the Holocene are scarce in the northeastern African region, one of the major candidates for the origin of Homo sapiens 76 and a gateway to the "out of Africa" migration of our species during the late Pleistocene. The 77 78 influence of climate on the dispersal of Anatomically Modern Humans from northeastern Africa particularly during the period ~120 to ~50 ka has been a subject of intense discussion 79 80 (e.g., Tierney et al., 2017 and references therein; Lamb et al., 2018). Recent discovery of Homo sapiens fossils dated to 177 to 194 ka in the Misliya cave in Israel (Hershkovitz et al., 81 2018) indicates that the "out of Africa" migration episodes have started earlier than the 82 83 previously thought period of migration (~120-50 ka). Discussions on influence of climate on human dispersal often rely on marine climate records from the Indian Ocean and 84 Mediterranean Sea. The recently published Lake Tana record from the northwestern 85 86 Ethiopian highland, largely covering the last ~150 ka (Lamb et al., 2018) is the only continental record available. In this paper, we present a high-resolution continental climate 87 record from an Ethiopian stalagmite (GM1) that grew intermittently around the last 88 89 interglacial, which is very pertinent to this discussion. Though it is not a continuous record over the whole period of the last interglacial, the growth phases of GM1 are dated at 90 particularly important periods of the MIS 5. The GM1 record, the oldest high-resolution 91 92 climate record so far published from Ethiopia and continental eastern Africa, is therefore very significant in an area where any kind of reliable continental climate records from this period 93 94 are scarce.

These annual-resolution records of δ^{18} O, δ^{13} C, trace elements and growth rate are from the Goda Mea Cave in Ethiopia (Fig. 1b). A combination of U-Th dates and lamina counting are used to identify the timing of the growth phases. Samples milled at annual resolution were analysed for δ^{18} O and δ^{13} C, and at decadal resolution for trace elements. Variogram, autocorrelation and spectral analyses of the geochemical and growth rate time series are used

to develop a conceptual model for the hydrology of the waters feeding the speleothem. The time series of δ^{18} O, δ^{13} C, trace elements and annual growth rate are then interpreted, with comparison to the published modern and Holocene stalagmites from the region and globally. Such high resolution, multi-proxy approach has been proved useful in reconstructing annual, in some cases seasonal, rainfall (e.g., Johnson et al., 2006).

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106 **2.** Methodology

107 *2.1. Site Description*

Goda Mea Cave was explored and surveyed in 2007 and a full description can be found in 108 Gunn et al. (2009). The cave is entered from a collapse doline and after about 30 m there is a 109 110 large flowstone deposit that almost fills the passage. A crawl beneath opens into a NE-SW oriented rift passage that is initially some 5 m wide by 1 m high but increases downstream to 111 10-15 m by 7 m. The cave ends in a 90 m x 40 m x 20 m high chamber formed by upwards 112 stoping as evidenced by abundant breakdown. Above the chamber there is ~25 m of sandy 113 limestone intercalated with some thin marl and mudstone layers towards the top, overlain by 114 115 a ~20 m thick calcareous sandstone, silt, carbonate rich shale and marl intercalation, which extends to ~1 m of soil at the surface. The limestone and sandstone-shale-marl units above 116 the cave form a continuous hydrogeological unit, connected by network of fractures. 117 118 There are numerous speleothems in the chamber including stalactites, stalagmites and a central column that is over 10 m high and 6 m diameter. The speleothems are mostly relict 119 with some evidence of re-solution and many of the stalagmites are fractured, most likely by 120 tectonic activity (Fig. 1c and d). Speleothem growth is generally focused along some aligned 121 zones below major fracture systems/brecciated fault traces crossing the hydrogeological unit 122 all the way up to the surface. 123

124	Some modern monitoring data (e.g., drip water chemistry) for the cave is presented in
125	Asrat et al. (2008). Drip water Ca ²⁺ and Mg ²⁺ concentrations in all analysed drip water
126	samples from this cave are 2.57 ± 0.65 mmol/L and 1.54 ± 1.12 mmol/L, respectively. The drip
127	water Ca^{2+} concentration in this cave is high as compared to the range of Ca^{2+} in the Mechara
128	caves (2.63±2.36 mmol/L) and falls within the range of values expected for "open system"
129	evolution (Baker et al., 2016). The high Ca^{2+} concentration can be attributed to the open
130	system calcareous sandstone/shale, marl and limestone hydrogeological unit, with the calcite-
131	cemented sandstones, carbonate rich shale, marl and limestone all contributing Ca^{2+} ions to
132	the drip waters.

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2.2. Sample description

GM1 is a large broken stalagmite found in the cave chamber. The 591 mm long 135 stalagmite was sectioned into two halves, and one half polished for lamina counting (Fig. 2). 136 The polished half shows continuous laminations, alternating between dense and porous/white 137 calcite, as well as visually recognizable growth hiatuses, marked by shifts in growth axis and 138 the stalagmite morphology. The other half of GM1 was milled using a hand held dental driller 139 for oxygen and carbon isotopes at ~0.6 mm resolution (966 samples), and trace element 140 analysis at ~5.5 mm resolution (103 samples). 38 samples for U-Th dating were similarly 141 drilled using a dental driller, with samples located either side of possible growth hiatuses, and 142 143 regularly spaced within growth phases.

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2.3. Geochemical analyses

146 U-Th analyses were undertaken by ICP-MS at the University of Melbourne, Australia, 147 following the method of Hellstrom (2003). Samples were dissolved in concentrated HNO₃ 148 and equilibrated with a mixed 229 Th $^{-233}$ U $^{-236}$ U tracer. U and Th were extracted in a single solution using Eichrom TRU resin before introduction to a Nu Plasma multi-collector ICP-MS, where isotope ratios of both elements were measured simultaneously. The decay constants of Cheng et al. (2013) were used, and detritally-corrected ages calculated using eqn. 1 of Hellstrom (2006) with an assumed initial [230 Th/ 232 Th] of 1.5 ± 1.5. Age-depth modelling combined floating annual laminae chronologies and U-Th analyses were as described in section 3.1.

155 Oxygen and carbon isotopes were analysed at the Stable Isotope Laboratory (SILLA), 156 University of Birmingham, UK. The calcite samples were reacted with phosphoric acid and 157 analysed using an Isoprime continuous flow mass spectrometer. By comparison with a 158 laboratory marble standard, the sample ¹⁸O/¹⁶O and ¹³C/¹²C ratios are reported as δ^{18} O and 159 δ^{13} C values in per mil (‰) versus VPDB. Analytical precisions are 0.07‰ for δ^{18} O and 160 0.04‰ for δ^{13} C on the standard marble (KCM).

Trace elements powders were analysed at University of New South Wales, Sydney.
Samples of approximately 0.05 g each were weighed directly into polypropylene vials. One
mL of 1-1 hydrochloric acid was added to each vial. The samples were sonicated for 15
minutes to ensure complete dissolution. The solutions were diluted to 10.0 mL with ASTM®
Type I water (Millipore® filtration system, Millipore® Corporation, Billerica, Massachusetts,
USA).

Diluted samples were analysed for Ca (317.933 nm) and Mg (285.213 nm) using the
PerkinElmer Optima[™] 7300DV ICP-OES (PerkinElmer, Shelton, USA). Ba and Sr were
analysed by PerkinElmer NexION 300D ICP-MS (PerkinElmer, Shelton, USA). Both
instruments were coupled with an ESI SC4 FAST sample introduction system (Elemental
Scientific, Inc., Omaha, USA) to minimise sample carryover.
The ICP-OES and ICP-MS were calibrated using certified multi-element standards in a

173 matrix of 2% HCl. Wavelength and analytical mass selection took into consideration spectral

interferences as well as sensitivities. Internal standards were added on-line via injection valve
to correct for physical interferences. Quality control check standards were run at selected
intervals in an unattended automatic analysis run, to ensure that the instrument performance
remained consistent over the length of analysis.

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2.4. Time series analysis

Statistical analysis on the annual growth rate time series followed the approach of 180 Mariethoz et al. (2012), which included the analysis of the first derivative of growth rate 181 182 (growth acceleration) to determine the flickering parameter (f), which is the magnitude of the anti-correlation at lag 1. Flickering ranges between -0.5 and 0, the more negative f values 183 indicating stronger flickering. Negative values of f are indicative of a karst store filling and 184 185 draining, as opposed to a climate forcing, and helps identify climatically sensitive speleothems. In addition, variogram analysis of the growth rate time series permits the 186 derivation of the information content (IC) and range (r) in the growth rate data, which helps 187 identify the signal: noise ratio in the data and the time over which useful information might 188 be expected. Stable isotope and annual growth rate time series data were also analysed for 189 their autocorrelation and spectral properties. As the data was evenly spaced in time, spectral 190 analysis was performed using discrete Fourier transforms, using the FFTW library within 191 192 Microcal Origin. Five windows were used (Bartlett, Hanning, Rectangular, Welsh and 193 Triangular) in order to investigate the extent of signal leakage.

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195 **3. Results**

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197 *3.1. GM-1 Chronology*

198 *3.1.1. Lamina*

199 Lamina were counted (in duplicate) and a total of 1356 lamina were identified with a 200 mean lamina thickness of 0.44±0.14 mm (ranging between 0.19 and 1.12 mm). This lamina thickness compares well to those reported for stalagmites in previous studies in the region 201 202 $(Ach-1, mean = 0.53 \pm 0.26 \text{ mm}, Bero-1 = 0.45 \pm 0.23 \text{ mm}, Asfa-3 = 0.32 \pm 0.11 \text{ mm}; Merc-1 = 0.45 \pm 0.23 \text{ mm}, Asfa-3 = 0.32 \pm 0.11 \text{ mm}; Merc-1 = 0.45 \pm 0.23 \text{ mm}, Asfa-3 = 0.32 \pm 0.11 \text{ mm}; Merc-1 = 0.45 \pm 0.23 \text{ mm}, Asfa-3 = 0.32 \pm 0.11 \text{ mm}; Merc-1 = 0.45 \pm 0.23 \text{ mm}, Asfa-3 = 0.32 \pm 0.11 \text{ mm}; Merc-1 = 0.45 \pm 0.23 \text{ mm}, Asfa-3 = 0.32 \pm 0.11 \text{ mm}; Merc-1 = 0.45 \pm 0.23 \text{ mm}, Asfa-3 = 0.32 \pm 0.11 \text{ mm}; Merc-1 = 0.45 \pm 0.23 \text{ mm}, Asfa-3 = 0.32 \pm 0.11 \text{ mm}; Merc-1 = 0.45 \pm 0.23 \text{ mm}, Asfa-3 = 0.32 \pm 0.11 \text{ mm}; Merc-1 = 0.45 \pm 0.23 \text{ mm}, Asfa-3 = 0.32 \pm 0.11 \text{ mm}; Merc-1 = 0.45 \pm 0.23 \text{ mm}, Asfa-3 = 0.32 \pm 0.11 \text{ mm}; Merc-1 = 0.45 \pm 0.23 \text{ mm}, Asfa-3 = 0.32 \pm 0.11 \text{ mm}; Merc-1 = 0.45 \pm 0.23 \text{ mm}, Asfa-3 = 0.32 \pm 0.11 \text{ mm}; Merc-1 = 0.45 \pm 0.23 \text{ mm}, Asfa-3 = 0.32 \pm 0.11 \text{ mm}; Merc-1 = 0.45 \pm 0.23 \text{ mm}, Asfa-3 = 0.32 \pm 0.11 \text{ mm}; Merc-1 = 0.45 \pm 0.23 \text{ mm}, Asfa-3 = 0.32 \pm 0.11 \text{ mm}; Merc-1 = 0.45 \pm 0.23 \text{ mm}, Asfa-3 = 0.32 \pm 0.11 \text{ mm}; Merc-1 = 0.45 \pm 0.23 \text{ mm}, Asfa-3 = 0.32 \pm 0.11 \text{ mm}; Merc-1 = 0.45 \pm 0.23 \text{ mm}, Asfa-3 = 0.32 \pm 0.11 \text{ mm}; Merc-1 = 0.45 \pm 0.23 \text{ mm}, Asfa-3 = 0.32 \pm 0.11 \text{ mm}; Merc-1 = 0.45 \pm 0.23 \text{ mm}, Asfa-3 = 0.32 \pm 0.11 \text{ mm}; Merc-1 = 0.45 \pm 0.23 \text{ mm}, Asfa-3 = 0.32 \pm 0.11 \text{ mm}; Merc-1 = 0.45 \pm 0.23 \text{ mm}, Asfa-3 = 0.32 \pm 0.11 \text{ mm}; Merc-1 = 0.45 \pm 0.23 \text{ mm}, Asfa-3 = 0.32 \pm 0.11 \text{ mm}; Merc-1 = 0.45 \pm 0.23 \text{ mm}, Asfa-3 = 0.32 \pm 0.11 \text{ mm}; Merc-1 = 0.45 \pm 0.23 \text{ mm}, Asfa-3 = 0.32 \pm 0.32 \pm 0.23 \text{ mm}; Merc-1 = 0.45 \pm 0.23 \text{ mm}; M$ 0.29±0.04 mm; Asrat et al., 2007; Baker et al., 2007; Baker et al., 2010). In these stalagmites, 203 the visible laminae have been demonstrated to be annual by comparison to the radiometric 204 dates. The GM1 laminae are similar in their appearance and thickness to these other 205 stalagmites. Examination of thick sections of GM1 at various levels of the growth phases 206 207 (Fig. 2) show continuous and regular visible laminae with alternating brownish calcite (Dark Compact Laminae, DCL) and thinner white calcite (White Porous Laminae, WPL) as defined 208 209 by Genty & Quinif (1996) and Genty et al. (1997). The presence of fine sediments on the 210 white porous calcite, and some dissolution features at the top of the DCL, suggests some seasonal infiltration variability (cf. Borsato et al., 2007). Overall, the regularly alternating 211 DCL/WPL laminae sequence, even without more obvious structures from infiltration 212 variability, indicate deposition under a seasonal hydroclimate regime (e.g. changes in drip 213 water supersaturation or cave air CO₂ concentration), where recharge was sufficient to 214 215 maintain continuous dripping to the stalagmite.

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217 *3.1.2. Growth hiatuses*

There are three major growth hiatuses based on the U-Th chronology (see below), and other possible minor growth hiatuses have been recognized by changes in the growth axis within the growth phases (Fig. 2). Visual examination of the three major growth hiatuses on the polished stalagmite and the thick sections show that the hiatuses between the four major growth phases are all marked by accumulation of fine detritus and brownish material on the top 2 mm sections, with no indications of dissolution. Such textural features are typical of theceasing of growth due to cessation of the drip source (Railsback et al., 2013).

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3.1.3. Stalagmite morphology

The morphology of the stalagmite changes from candle stick shaped, with regular nearly 227 horizontal lamina on and off the growth axis for most of the first growth phase to upwards-228 thinning, laterally less extensive layers with laminae rapidly changing to sub-vertical angle, 229 off the growth axis, just below the first hiatus (Fig. 2). The second growth phase above the 230 231 first hiatus then gets broader at its axis with rapid flowing/dripping down the sides of the stalagmite forming nearly vertical lamina. The third growth phase shows similar morphology 232 to the second though it rapidly thins towards the top below the third major hiatus. The last 233 234 growth phase has relatively broader shape with significant deposition along its axis. The morphology of the stalagmite changing with the hiatus position is a clear demonstration of 235 the changing amount and concentration of calcite in the dripping water. It shows a general 236 237 drying out of the drip source towards the tops of the three older growth phases, while the last growth phase is marked by an increased drip rate throughout the growth period. 238

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3.1.4. U-Th dates and annual growth rate

The 38 U-Th dates on the sample (Fig. 3, Table 1) demonstrate 4 periods of growth and confirm the presence of 3 hiatuses. Several age inversions are present, and one short growth phase containing 37 laminae was undated.

The U-Th ages were used to constrain a chronology based on the annual laminae. Firstly, the longest phase of speleothem growth (from ~127 mm from the top, to the base at 591 mm) contained 29 very similar U-Th ages (a mean and standard deviation of 130.0 ± 3.5 ka, with an average uncertainty on individual analyses of 1.3 ka), providing further evidence that the 248 1087 laminae present in this growth phase are likely to be annual. Secondly, following 249 established methods (Asrat et al., 2007; Dominguez-Villars et al., 2012; Baker et al., 2015), we compared two approaches to tie the lamina chronology to the U-Th ages. The first 250 251 approach was as follows: within each growth phase, each U-Th age was adjusted by using its relative laminae age to obtain an equivalent U-Th age for the date of the start of each growth 252 phase. Taking the mean and standard deviation, this yielded growth phases starting at 253 129.3±2.7 ka, 120.7±1.7 ka, and 108.3±0.2 ka. We compared this approach to that calculated 254 from linear regression applied to conventional age-depth plots. In this case, we used only the 255 U-Th ages with a $[^{230}\text{Th}/^{232}\text{Th}]$ >1000, presuming they would be the most accurate. This 256 approach yielded a date for the start of deposition for two of the four growth phases of 257 258 129.2±1.7 ka and 120.6±0.3 ka. The two approaches therefore give consistent dates for the 259 start of deposition that agree with the analytical error of individual analyses. GM1 deposition periods are therefore ascribed to four phases: 1087 years commencing 260 129.3±2.7 ka, 54 years of deposition at 120.7±1.7 ka, a 37-yr long undated growth phase, and 261 176 years of deposition from 108.4±0.3 ka. The ~129.3 ka growth phase occurs, within 262 dating uncertainty, at Termination II or the early part of the last interglacial (Cheng et al., 263 2009), and the 120.7 ka deposition immediately post-dates the full interglacial. The growth 264 phase at 108.5-108.3 ka falls within the isotope stage 5c interglacial. 265 The annual growth rate for GM1, determined from the annual lamina thickness, is 266 267 presented in Figure 4. Mean growth rate does not vary between growth phases: 0.43 ± 0.14 mm/yr (129.3 ka), 0.41±0.17 mm/yr (120.7 ka), 0.53±0.15 (undated), and 0.47±0.14 mm/yr 268 (108.4 ka). 269

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271 *3.2. Oxygen and carbon isotopes*

The oxygen and carbon isotope (δ^{18} O and δ^{13} C) data were ascribed to an annual lamina 272 and are presented in Figures 4. Scatter plots of δ^{18} O and δ^{13} C, including the analysis of the 273 isotopes along a lamina (the 'Hendy test') are shown in Figure S1. 274 δ^{18} O varies significantly between growth phases (-7.77±0.57‰ (129.3 ka), -275 3.82±0.61‰ (120.7 ka), -6.05±0.76‰ (undated), and -6.31±0.59‰ (108.4 ka). Within each 276 growth phase, δ^{18} O can be described as having long periods of relatively invariant 277 278 composition (e.g., ±0.33‰ for the first 1000 years of the 129.3 ka deposition period, and $\pm 0.23\%$ in the first 150 years of the 108.4 ka deposition period), as well as periods of rapid 279 change. For example, δ^{18} O increased from -7.8% to -5.3% in four years (and to -4.5% after 280 13 years) at the end of the 129.3 ka growth phase, and from -6.2% to -4.4% in nine years at 281 282 the end of the 108.4 ka growth period.

 δ^{13} C is characterised by low inter-sample variability, with the presence of long-term trends. For example, in the 129.3 ka growth phase, the standard deviation of δ^{13} C over any 50-year period is between 0.1‰ and 0.3‰, but over the whole 1087 years of deposition, δ^{13} C trends from –1‰ to –4‰. This 3‰ change in δ^{13} C with the growth phase is as great as the variability between growth phases.

Figure S1 shows the relationship between δ^{18} O and δ^{13} C, both throughout the time series 288 as well as along growth layers ('Hendy tests'). The Hendy tests suggest a 1‰ increase or 289 decrease in isotope composition is possible along a growth layer, which is greater than the 290 inter-annual variability of δ^{13} C and δ^{18} O. For stalagmite GM1, there is no evidence for near-291 equilibrium deposition: modern and Holocene stalagmites demonstrate isotope fractionation 292 of 1-2‰ (Asrat et al., 2007; Baker et al., 2007; 2010), and similar deposition conditions 293 appear to apply to stalagmite GM1. Based on these works, we have quantified these 294 fractionation processes and confirmed that they operate in the same direction as the climate 295 296 forcing, which has also been observed by other works (e.g., Dorale and Liu, 2009).

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3.3. Trace elements

Ba/Ca, Mg/Ca and Sr/Ca show similar patterns to δ^{13} C and δ^{18} O, although sampled at a 299 lower temporal resolution. A long-term trend to lower ratios in the 129.3 ka growth phase 300 matches that observed in δ^{13} C. Significant short-term changes in trace element composition 301 occurs at the same time as the increases in δ^{18} O and δ^{13} C at the end of the 129.3 ka growth 302 phase, and within the 120.7 ka growth period. At the end of the 129.3 ka growth phase, a 303 change in gradient of Mg/Ca, Sr/Ca and Ba/Ca lasted for ~170 years, indicative of a drying 304 trend. This was followed by the δ^{18} O increase of 2.2‰ that occurred over 4 years, and then 305 306 an increased variability in Sr/Ca, Mg/Ca and Ba/Ca (Mg/Ca increases, Mg/Ca and Ba/Ca decrease) until growth stops 28 years later. In contrast, the trace element response at the time 307 of a 2‰ increase in δ^{18} O at the end of the 108.4 ka growth phase is muted and trends to lower 308 values. The greatest range in trace elements occurs in the 120.7 ka growth phase, where Sr/Ca 309 310 increases, and Mg/Ca and Ba/Ca have opposing increasing and decreasing trends.

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312 *3.4. Time series analysis on stable isotope and growth rate time-series*

Following Mariethoz et al. (2012), and as described in section 2.4, the growth rate time series variogram properties was investigated. Due to the short duration of several of the growth phases, only the longest time series at 129.3 ka was analysed. The results are plotted in Figure 5a, where variogram analysis on stalagmite GM1 is compared to previous published stalagmite statistics on growth rate series. The autocorrelation and spectral properties of the growth rate, δ^{18} O and δ^{13} C series were also investigated (Figs 5b and c).

Stalagmite GM1 growth at 129.3 ka has evidence of 'flickering' (f=-0.33), that is a

320 growth acceleration that flickers around a mean value (Mariethoz et al., 2012). 'Flickering'

321 has been explained as growth rate sensitivity to the filling and draining of a karst store, which

is trying to reach a dynamic equilibrium, with reported values between -0.24 (low flickering,
potential climate signal) and -0.39 (high flickering, potential karst hydrology signal).

Variogram analysis shows that the stalagmite has low range (r, 20.5 years) in the growth 324 325 rate record. Growth rate therefore has no 'memory' of previous growth rates longer than this timescale, indicative that the karst store(s) that feed the stalagmite are relatively small. The 326 GM1 range is the lowest reported to date for an annually laminated speleothem. The 327 Information Content (IC) of growth rate, which is the balance of the signal in the variogram 328 and the noise, is 56%, indicating that the stalagmite growth rate record contains a greater 329 330 proportion of signal than noise. In comparison to previously published records (Fig. 5a), stalagmite growth rate statistical properties lie in region A, with stalagmites that have a high 331 information content, relatively low flickering and range, and where growth rate has proven 332 333 useful in paleoclimate reconstruction.

Autocorrelation of growth rate, δ^{13} C and δ^{18} O time series are presented in Figure 5b for 334 the three longest growth phases (~129, ~120 and ~108 ka). Significant autocorrelation can be 335 observed for δ^{13} C and δ^{18} O for the ~129 and ~108 ka growth phases, with autocorrelation 336 >0.6 at 15-year lag. Autocorrelation for δ^{13} C is stronger than for δ^{18} O, indicative of 337 additional smoothing of the δ^{13} C, likely from the soil carbon store. Growth rate has very low 338 339 autocorrelation (<0.4 after 4 years lag), in agreement with the observed flickering of growth rate. The ~120 ka growth phase has lower autocorrelation of δ^{18} O and δ^{13} C than the other two 340 growth phases, suggesting limited mixing or smoothing of these proxies during this short 341 growth phase. Low autocorrelation would agree with the observed highest variability in trace 342 343 elements at this time.

Spectral analysis was also undertaken (Fig. 5c) for the longest continuous growth phase at ~129 ka. Bartlett, Hanning, Rectangular, Welsh and Triangular windows were used to explore the spectral properties for δ^{18} O and growth rate time series. Growth rate has a 17-18

year peak that is not statistically significant at 95% confidence, and two other peaks (31-33 and 53-59 years), which are longer than the range (r, 20.5 years) and are likely to be harmonics of the 17-18 year frequency. δ^{18} O also has only weak and insignificant spectral power, not surprising given the low variability in the δ^{18} O data.

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352 **4. Discussion**

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354 *4.1. Holocene stalagmite records*

Previous cave research in Ethiopia has included limited cave drip water and climate 355 monitoring during sampling expeditions to the Mechara region of Ethiopia between 2004 and 356 2008 (Asrat et al., 2008) (Fig. 1), and the analysis of modern and Holocene stalagmite 357 samples. Modern calibration studies of stalagmite δ^{18} O of carbonate shows evidence of 358 climate sensitivity, despite deposition out-of-equilibrium (Baker et al. 2007; 2010). The latter 359 is potentially due to both rapid degassing and evaporation. δ^{18} O and growth rate correlations 360 with climate are sample-specific. Drip-specific flow-paths determine whether a stalagmite 361 has a proxy which is sensitive to the 'big' rains, or to the relative amount of rain in the 'big' 362 and 'small' rain seasons, or neither. For example, modern stalagmite δ^{18} O and growth rate 363 records were reported from two stalagmites from Rukiesa Cave (Baker et al., 2007). The 364 annual nature of the laminae was confirmed by ¹⁴C analyses and comparison to the modern 365 atmospheric bomb carbon peak. In these samples δ^{18} O and growth rate were shown to have a 366 correlation with the ratio of 'small' to 'big' rainfall and total summer rainfall, respectively. A 367 sample specific climate sensitivity of δ^{18} O and growth rate was observed, which probably 368 reflects the karst hydrogeology and its effect on individual water flow-paths. A loss of 369 370 climate correlation was also observed in one sample during a period of high growth rates.

371	δ^{18} O and growth rate time-series both exhibit multi-decadal variability in two stalagmite
372	samples deposited in the Holocene. Such variability may be an amplification of extremes in
373	hydroclimate (e.g., drought years) or rainfall isotopic composition, due to the non-linear
374	replenishment or drainage of karst stores (Baker et al., 2013). A mid-Holocene record from
375	Achere Cave (sample Ach-1) had laminae that were demonstrated to be annual by
376	comparison to U-Th dates (Asrat et al., 2007). In this stalagmite a 18-21 yr periodicity in
377	growth rate and δ^{18} O occur. δ^{18} O has a greater variability than δ^{13} C, indicative of the
378	variability being driven by variations in the extent of evaporative enrichment of δ^{18} O. A
379	discontinuous Holocene record from Bero Cave had six growth phases over the last 7800
380	years (Baker et al., 2010). Mean stalagmite δ^{18} O is 1.2‰ higher than that predicted by
381	forward modelling, and a multi-decadal variability in $\delta^{18}O$ and growth rate was again
382	observed (Baker et al., 2010). δ^{18} O from this stalagmite was indicative of both rapid
383	degassing and the additional enrichment, probably due to evaporation.
384	Stalagmite growth phases are relatively short (10^3 - 10^4 years) due to the tectonically active
385	nature of the region, which can change water flow paths. This is observed in stalagmites from
386	all three caves, and has been explained by changes in flow regime or to the relative position
387	of a growing stalagmite caused by tectonic activity related to the East African Rift (Asrat,
388	2012). Physically anomalous laminae within an otherwise regular and visible annual laminae
389	sequence, frequent deviations from vertical growth axis, and abrupt changes in stalagmite
390	morphology, as well as the tectonically-controlled formation of the larger cave system,
391	further confirm the influence of tectonics and recorded earthquakes in the region (see Fig. 1a)
392	on the length of the growth phases (Asrat et al., 2008; Asrat, 2012).
393	

4.2.Conceptual growth model

The combination of stable isotope and trace element geochemistry, growth rate, statistical analyses, and observations of the laminae types and stalagmite shape, allow us to build a conceptual model for stalagmite GM1 (Fig. 6).

398 Firstly, the stalagmite probably had continuous deposition for more than 1000 years, and during this period isotope and trace element composition has low variability and high 399 autocorrelation. This homogeneity of δ^{13} C, δ^{18} O and trace elements suggests a drip water 400 source which is well mixed, enough to obscure any annual to decadal scale variability in δ^{18} O 401 and maintain dripping. The continuous deposition of the laminae and the candle-stick shape 402 of the stalagmite before it narrows down towards the tip of this growth phase (the last few 403 years of growth) supports a continuous drip source. We propose this water comes from 404 matrix flow of the porous sandstone and sandy limestone, which was channelled to the drip 405 406 source by a network of small fractures. The annual laminae are driven by this flow regime, which provides the necessary seasonal variability in drip water hydro-geochemistry. 407 Combined with the evidence from flickering, we infer the variations in lamina thickness are 408 409 driven by the karst hydrology and not by the cave environment. In the ~129 ka growth phase, 410 these would have to maintain high levels of saturation for the initial ~1000 years of deposition. 411

412 Secondly, the four growth phases of GM1 reflect the changing karst hydrologic regime above the cave. The ~129 ka growth phase is marked by the dominance of a continuous 413 414 supply of water from 'matrix' flow for most of its growth period, which rapidly dried out a few tens of years before the hiatus. In contrast, the ~120 ka growth phase reflects a rapid 415 'fracture' flow, following a possible tectonic event, which did not maintain growth for a long 416 417 time before it abruptly shut off as marked by the rapid increase in trace element ratios and δ^{18} O. The third, undated growth phase shows similar features to that of the second, only the 418 growth period was shorter suggesting a rapid start to dripping and subsequent exhaustion 419

420 from a 'fracture' source, as marked by the nearly vertical laminae down the sides of the 421 stalagmite and the abrupt narrowing towards its top. The ~108 ka growth phase is again 422 marked by a more 'porous' matrix flow which maintained growth for longer period, attested 423 by the regular laminae, broad stalagmite shape, generally low variability trace element ratios 424 and depleted δ^{18} O.

Thirdly, growth rate varies annually, shows evidence of flickering, and has a range of ~ 20 425 years and a spectral peak at 17-18 years. Given the relative homogeneity of the stable isotope 426 427 and trace element signals, the growth rate variability has to occur subsequent to the mixing of the water. Dissolutionally-enlarged fractures or a network of small conduits would allow 428 limited water storage, permitting degassing of karst water and prior calcite precipitation 429 430 (PCP) as well as drip rate variations, both affecting growth rate variability. We conceive this store to have a proportional volume of approximately 20 years of recharge (see later). In 431 drying conditions, water from this store could maintain dripping and deposition for short time 432 periods. 433

Fourth, a pre-existing fracture zone/brecciated fault trace which might have been 434 435 reactivated during subsequent tectonic activities, extending from the surface, through to the store and the cave roof, permits fracture flow to stalagmite GM1. This would permit short-436 duration recharge, probably after high magnitude / frequency rainfall events, in the absence 437 438 of saturated porous sandstone and limestone aquifer. This would explain the short duration, high geochemical variability, ~120 ka growth phase. The absence of 'stored' water and 439 subsequently rapid exhaustion of the fracture flow/drip source is supported by the nearly 440 vertical lamina depositing down the sides of the stalagmite, high variability/rapid increase in 441 the δ^{18} O and trace element ratios and low autocorrelation in δ^{13} C and δ^{18} O. 442

443 The varying trends between the δ^{18} O and δ^{13} C, and trace element ratios from one growth 444 phase to the other suggests that a single kinetic fractionation process does not dominate our

proxy records, rather we infer a hydrological control based on climatic and tectonicprocesses.

Our conceptual model explains other features of GM1 geochemistry and growth rate. The 447 geochemistry at the end of the 129.3 ka growth phase can be interpreted as a decline in the 448 saturation or water level in the porous media, leading to an increase in δ^{18} O and trace 449 elements as dripping is maintained just from the smaller store. Before growth stops, a change 450 in gradient of Mg/Ca, Sr/Ca and Ba/Ca indicates a drying trend which lasts for ~170 years, 451 followed by a δ^{18} O increase of 2.2‰ that occurred over 4 years, and then increased 452 variability in Sr/Ca, Mg/Ca and Ba/Ca until deposition stops 28 years later. Many of the 453 previously studied stalagmites from Ethiopia such as Ach-1 (Asrat et al., 2007) show similar 454 features, which could be attributed to the specific geological setting of the region where 455 456 earthquake/tectonics play a strong role in shifting the relative position or the extent of the major 'fracture' flow routes for such short-phased growths, leading to growth maintained for 457 short time longer from the smaller 'matrix' flow. 458

459 The similarity in values for the range r (20.5 yrs), the spectral frequency f (17-18 yrs), 460 and the observation that it takes 24 yrs for the stalagmite to stop growing, all suggest the presence of a water store that can hold ~20 years of recharged water. The multi-decadal 461 462 growth rate frequency of 17-18 yrs, although insignificant, is in agreement with that observed from Holocene stalagmites in the region (Ach-1, Bero-1; Baker et al., 2010), and similar to 463 observed variability in the modern rainfall pattern and subsequent flow at the upper Blue Nile 464 (Taye and Willems, 2012). Plausible climatic forcing over this timescale includes changes in 465 Atlantic and Indian Ocean sea surface temperature and variability in the movement and 466 467 intensity of the ITCZ and its effect on Ethiopian rainfall (Degefu et al., 2017). However, the similarity of f and r suggests that any climate forcing in GM1 growth rate variability may be 468

469 amplified by the size of karst water store, or be karstic rather than climatic, the latter something previously observed in forward modelling studies (Baker et al., 2013). 470 Finally, the shift to higher isotope values within a growth phase can be explained by our 471 472 conceptual model as a change from porous flow being the dominant water source to a dominance of fracture flow. It suggests that the observed 2‰ shift could be indicative of a 473 water that has undergone additional kinetic or evaporative isotope fractionation. Both 474 475 fractionation processes had been previously inferred as occurring in both Modern and Holocene Ethiopian speleothems, and from 'Hendy tests' on GM1, to a similar extent (up to 476 1‰). The implication for the climatic interpretation of stalagmite δ^{18} O is that variability of 477 up to 2‰ cannot be directly ascribed to climatic forcing, but larger changes cannot be 478 explained by fractionation processes. Similar rapid shifts in δ^{18} O of +2‰ within a period of 6 479 years have been identified in the Hulu cave speleothems (Treble et al., 2007). 480

481

482 *4.2. The climate record*

We can compare the δ¹⁸O composition for each GM1 growth phase with published
Holocene stalagmite data from Ethiopia (Asrat et al., 2007; Baker et al., 2007; 2010), as well

as the modelled solar insolation for 15°N (Laskar et al., 2004), and other archived speleothem

486 δ^{18} O records along the monsoon path and the "downstream" countries (China, Cheng et al.,

487 2017; Israel, Bar Matthews et al., 1999; 2003). This comparison is shown in Figure 7. In GM-

488 1 the δ^{18} O composition (-7.77±0.57‰ (129.3 ka), -3.82±0.61‰ (120.7 ka), -6.05±0.76‰

(undated), and $-6.31\pm0.59\%$ (108.4 ka)) is generally isotopically more negative compared to

- 490 both modern (Merc-1: -1.22±0.31‰; Asfa-3: -1.37±0.37‰; Baker et al. 2007), and
- 491 Holocene (Bero-1: -3.42±1.45‰, Baker et al., 2010; Ach-1: -3.20±0.35‰, Asrat et al.,
- 492 2007) samples from the region. Even allowing for kinetic fractionation and non-equilibrium
- deposition of up to 2‰ in all samples, GM1 δ^{18} O composition at 129.3 ka, 108.4 ka, and an

undated growth phase, is more negative than any Holocene stalagmites from the region.
Combined with our conceptual understanding of GM1 deposition, we can be certain that
these growth phases and lower isotope composition are indicative of wetter conditions and
sustained recharge.

The GM1 record is the first high-resolution last interglacial continental climate record 498 and among the few climate records of any resolution from Ethiopia so far published. A deep 499 seismic and near-continuous core record of the last 150,000 years from Lake Tana on the 500 501 Northwestern Ethiopian highlands used geochemical proxies (sediment Ca/Ti ratio) for 502 climate-driven lake level fluctuations (Lamb et al., 2018). The oldest cave sediment records from the Southeastern Ethiopian highlands goes back only to 63±7 ka (Tribolo et al., 2017). 503 504 The four phases of the GM1 record are dated at particularly important periods of the last 505 interglacial. Noting the quantified age uncertainties (see section 3.1.4), they provide high resolution snapshots from some critical time-windows. The two long growth phases at ~129 506 ka and ~108 ka, which we conceptualise as being dominated by a sustained porous/matrix 507 flow regime, match maximum summer insolation at 15°N. This suggests that though internal 508 growth variability may be dominated by karst hydrology above the cave, the GM1 growth as 509 a whole and the geochemical proxies were responding to climate forcing. 510

Comparison of the GM1 δ^{18} O record with the China composite δ^{18} O record (Hulu and 511 Dongge caves; Cheng et al., 2016), the Soreq cave (Israel) δ^{18} O record (Bar-Mathews et al., 512 1999; 2003) (Fig. 7) suggests a similar relationship for all three sites, with wet conditions 513 (lower δ^{18} O and peak summer insolation) during the ~108 ka growth period, and dry or 514 drying conditions (higher δ^{18} O and low summer insolation) at the ~120 ka growth period. 515 The Soreq cave δ^{18} O record from central Israel in the Levant has been shown to indicate 516 enhanced rainfall (Bar-Mathews et al., 1999) and could be a "downstream" indicator of a 517 stronger northeast African monsoon (Tierney et al., 1999). At ~129 ka, the lower δ^{18} O and 518

519 wetter conditions in Ethiopia occur at the summer insolation maxima, but may occur before the isotope response observed in Israel and China. Dating uncertainty in the GM1 record 520 prevents a more precise interpretation, but it does raise the possibility of the intensification of 521 522 the East African Monsoon before other northern hemisphere monsoon systems at the start of the last interglacial. The Lake Tana (Northwestern Ethiopian highlands) sediment Ca/Ti 523 record (Fig. 7) indicates an abrupt increase in moisture even earlier at ~132 ka, leading to 524 stable high lake level conditions during the period ~132 ka to ~95 ka, with some brief dry 525 episodes (Lamb et al., 2018). This period is defined by a generally flat trend of Ca/Ti, 526 truncated by very brief high Ca/Ti excursions. The major wet conditions during the ~129 ka 527 and 108 ka growth phases of GM1 are generally consistent with this predominantly high lake 528 level phase, although they do not particularly correspond to the lowest values of Ca/Ti. 529 530 As speleothem records of the Holocene from the region (Fleitmann et al., 2007) show, decreasing δ^{18} O values during the early Holocene indicate a rapid northward migration of the 531 summer ITCZ and intensification of the rain belt of the Indian Summer Monsoon. On the 532 other hand, the southward migration of the ITCZ during the middle to late Holocene, marked 533 by increasing δ^{18} O values in speleothems, led to weakening of the associated summer 534 monsoon. Similar studies from Madagascar (e.g., Voarintsoa et al., 2017) have also shown 535 the link between speleothem δ^{18} O values and the ITCZ migration. The ~129 ka growth phase 536 of GM1 could therefore be effectively considered as an annual record of relative ITCZ 537 position ~129 ka BP. 538

Ethiopia/northeast Africa has been considered as the origin of Anatomically Modern
Humans (White et al., 2003; McDougall et al., 2005) and possibly the region out of which
Anatomically Modern Humans dispersed during MIS 5-3 or during earlier migration episodes
(*cf.* Hershkovitz et al., 2018). Considering the significant debate about the role climate
variability played in human evolution and dispersal of Anatomically Modern Humans "Out of

Africa" (e.g., Tierney et al. 2017 and references therein; Lamb et al., 2018), high-resolution 544 speleothem records from Ethiopia and northeastern Africa such as those of GM1 can shed 545 light on this debate. The GM1 record for instance shows that the earliest human migration 546 during MIS 5, confirmed by the presence of ~110-80 ka old Anatomically Modern Human 547 fossils in Israel (Grün et al., 2005), occurred during a major wet event in northeast Africa. 548 The Tana Lake record confirms that the northwestern Ethiopian highlands experienced 549 relatively stable moist climate during MIS 5c-e (Lamb et al., 2018). This supports earlier 550 conclusions that human migration occurred during humid conditions, as such conditions 551 552 provided humans "green corridors" to overcome inhospitable deserts (e.g., Timmerman and Friedrich, 2016). However, the major episode of human migration occurred during 50-75 ka 553 (Nielsen et al., 2017), and marine records from the Gulf of Aden show that this migration 554 555 event occurred during a sustained dry condition in northeast Africa (Tierney et al., 2017), while the Lake Tana record shows more complex climate variability during this period (Lamb 556 et al., 2018). Future research is required on the speleothems from Goda Mea and Aynage 557 caves, some of which have been dated to the critical period of 120-50 ka. 558

559

560 **5.** Conclusions

Stalagmite GM1 was deposited discontinuously around the time of the last interglacial, at 561 ~129 ka, 120 ka, an undated growth phase, and ~108 ka. Variogram analysis of growth rate 562 shows a low range (20.5 years), some flickering (-0.33) and good information content (56%), 563 indicative of a stalagmite fed by a karst water store of limited volume. Oxygen and carbon 564 isotopes and trace elements generally have low variability, indicative of a second, well-mixed 565 566 water source feeding the stalagmite. Stalagmite GM1 provides a high-resolution insight into stalagmite hydrogeochemical responses to environmental change prior to growth hiatuses. 567 Multi-decadal variability of frequency 17-18 years, though statistically not significant at 95% 568

confidence, is present, but only in the growth rate time series, and is slightly less than the
range in the growth rate record. A climatic or karstic forcing of this spectral frequency cannot
be determined.

Our conceptual model for the stable isotope, trace element and growth rate records in 572 GM1 allows the interpretation of the stalagmite geochemical time series. Importantly, all 573 three proxies were necessary to adequately understand the processes forcing them, and 574 575 whether they contained a climatic or karstic signal. Only through this approach were we able to confirm that low δ^{18} O at ~129 ka and ~108 ka can be attributed to wetter climatic 576 conditions. These two growth phases occur at the same time as solar insolation maxima for 577 578 15°N, and suggest a direct solar forcing on rainfall in Ethiopia at these times, influencing the northward migration of the ITCZ and the associated rain belt of the Indian Summer 579 580 Monsoon, of potential relevance for early modern human migration out of the region. 581

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- 597 http://www1.ncdc.noaa.gov/pub/data/paleo/speleothem/asia/china/cheng2016composite.txt
- 598 ftp://ftp.ncdc.noaa.gov/pub/data/paleo/speleothem/israel/soreq_peqiin_2003.txt
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Table 1. U and Th isotope data and age determinations (in depth order) for stalagmite GM1.
Square brackets indicate activity ratios. Ages shown are corrected for an initial
[²³⁰Th/²³²Th] of 1.5 and a 100% uncertainty, which is incorporated into the age
uncertainty.

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731 Figure Captions

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Fig. 1. (a) Regional structural setting of Ethiopia showing the location of Mechara. Lake 733 734 Tana, and the epicentres of the major earthquakes in the Main Ethiopian Rift and the 735 adjoining highlands are marked (Note that earthquake epicentres in the northern Afar depression are not represented). Insets show the mean position of the ITCZ in July 736 (Boreal summer) and January (winter) over Africa; and the mean monthly rainfall (mm) 737 738 and mean monthly temperature of the Mechara region, at the Bedesa Meteorological Station (1994-2014 data from the Ethiopian Meteorological Agency). Location of (b) is 739 740 marked by a broken triangle around the location of Mechara (modified from Asrat et al., 2008); (b) The topography, geology, structure and drainage system of the Mechara karst 741 742 area and locations of the entrances to the caves mentioned in the text; (c) Goda Mea cave 743 (surveyed according to BCRA Grade 3 using tape, compass and clinometer); (d) a photograph of the main chamber of the Goda Mea cave interior showing the collapse 744 chamber on which grew several speleothems following a major fracture system 745 746 (photograph by J. Gunn). Figures (a) and (b) modified from Asrat et al., 2008; Fig. (c) modified from Gunn et al., 2009. 747

Fig. 2. GM1 hand-section in both scanned image (left) and sketch (middle), showing the four
major growth phases, locations of the major and minor growth hiatuses, and sampling for
isotopes, trace elements and U-Th analyses. Right: photomicrographs of thick sections
from across the major growth hiatuses showing a clear evidence of growth stoppage with
no apparent dissolution.

Fig. 3. U-Th data for stalagmite GM1. (a) Corrected U-Th ages vs depth for all analyses. (b)

754 238 U concentration vs depth (c) [230 Th/ 232 Th] vs depth (d) Initial [234 U/ 238 U] vs depth (e)

755 U-Th ages vs depth for samples with $[^{230}Th/^{232}Th] > 1000$. In all plots, the three major

hiatuses are shown as vertical dashed lines.

Fig. 4. GM1 times series for the geochemical proxies. From top: growth rate, δ^{13} C, δ^{18} O,

Sr/Ca, Mg/Ca, Ba, Ca. Note the axis breaks on the x-axis, which permit equal scaling ofdata on the time axis.

Fig. 5. (a) Scatterplot of variogram parameters range, information content and flickering for

the 129.3 ka growth phase; (b) Autocorrelation of growth rate, δ^{13} C and δ^{18} O time series

for the three U-Th dated growth phases; (c) Spectral analysis on the growth rate and δ^{18} O

times series for the 129.3 ka growth phase. The 17-18 years peak, though statistically not

significant is marked.

Fig. 6. Conceptual model for the deposition of stalagmite GM1.

Fig. 7. Comparison of climate proxy records. (a) Insolation at 15°N (Laskar et al., 2004); (b)

767 Chinese composite stalagmite δ^{18} O record (Cheng et al., 2016); (c) Soreq Cave δ^{18} O

record (Bar-Matthews et al., 1999; 2003); (d) Lake Tana sediment Ca/Ti record (Lamb et

al., 2018); (e) Ethiopian stalagmite δ^{18} O composite. Box-plots show median, inter-

quartile range and range for each stalagmite; shading represents different caves (grey –

771 Bero Cave; Green – Rukiesa Cave; Orange – Achere Cave; Cyan – Goda Mea Cave).

- Note the x-axis break; vertical shading aligns the Ethiopian records to the other timeseries.
- Fig. S1. Scatter plot of δ^{13} C vs δ^{18} O for all growth phases. 'Hendy tests' along growth
- laminae are shown in colour (Lines 1-7): their location is shown in Fig. 2 (HL1-HL7).
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